# The R-xion: Protecting the Peccei-Quinn Symmetry by means of Non-Anomalous Discrete *R* Symmetries.



#### Kai Schmitz

Kavli Institute for the Physics and Mathematics of the Universe (WPI)

Todai Institutes for Advanced Study, University of Tokyo, Kashiwa, Japan

Based on arXiv:1308.1227 [hep-ph].

In collaboration with Keisuke Harigaya, Masahiro Ibe and Tsutomu T. Yanagida.

SUSY, ICTP Trieste, Italy | Dark Matter and Cosmology | August 26, 2013

Minimal Extension of the MSSM with a Peccei-Quinn Symmetry

Phenomenological Constraints

Minimal Extension of the MSSM with a Peccei-Quinn Symmetry

Phenomenological Constraints

## The Peccei-Quinn Solution to the Strong CP Problem (I)

#### The strong *CP* problem:

- ► Axial QCD anomaly  $\to \mathscr{L}_{\text{QCD}}^{\text{eff}} \supset \bar{\theta} \frac{\alpha_s}{8\pi} \text{Tr} \big[ G_{\mu\nu} \widetilde{G}^{\mu\nu} \big]$  w/  $\bar{\theta} = \theta + \text{arg} \{ \det M_q \}$ .
- ▶ Neutron electric dipole moment  $d_n \simeq 5 \times 10^{-16} \,\bar{\theta} \, \mathrm{e \ cm} \lesssim 3 \times 10^{-26} \, \mathrm{e \ cm}$ .
- ▶ QCD vacuum angle  $\bar{\theta} \lesssim 10^{-10}$ . Expectation  $\bar{\theta} \sim \mathcal{O}(1)$ .  $\Rightarrow$  Why so tiny?

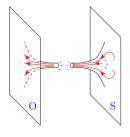
#### One solution: Promote $\bar{\theta}$ to dynamical real scalar field with VEV at 0.

[Peccei & Quinn '77; Weinberg '78; Wilczek '78]

- ▶ The axion: pseudo-NG boson of a spontaneously broken global  $U(1)_{PO}$ .
- QCD instanton-induced effective potential after the QCD phase transition:

$$V_a = \Lambda_{\mathrm{QCD}}^4 \left[ 1 - \cos \left( ar{\theta} - a/f_a 
ight) 
ight] \, , \quad \langle a 
angle = ar{\theta} \, f_a \, .$$

#### The Peccei-Quinn Solution to the Strong CP Problem (II)



rig. from J.E. Kim, 1308.0344 [nep-tn]

## But: Any global symmetry is believed to be broken by quantum gravity effects! [Kamionkowski & March-Russell '92: Barr & Seckel '92: Holman et al. '92: Banks & Seiberg '11]

- ▶ Why is  $\bar{\theta}$  so small? → Why is the global PQ symmetry of such high quality?
- Answer: Approximate accidental U(1)<sub>PO</sub> due to exact gauge symmetry.

Our idea: Protect PQ symmetry by means of gauged discrete R symmetry,  $Z_N^R$ .

## An Anomaly-Free Discrete R Symmetry for the MSSM (I)

#### Strong motivation for $Z_N^R$ in SUSY phenomenology and model building:

[Giudice & Masiero '88; Yanagida '97; Dine & Kehayias '10] [Dimopoulos & Georgi '81; Sakai & Yanagida '82; Weinberg '82] [Izawa & Yanagida '97]

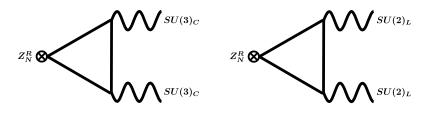
- ▶ No large  $\mu$  term, no dangerous proton decay, no large  $\langle W \rangle$  (i.e. negative  $\Lambda$ ).
- Possibly remanant subgroup of continuous stringy U(1)<sub>R</sub> symmetry.
- ▶ If  $U(1)_R$  gauged, remnant  $Z_N^R$  gauged and not broken by quantum gravity.

#### Rendering the $Z_N^R$ symmetry anomaly-free:

[Ibanez '93]

- ▶ Generation-independent  $Z_N^R$  with N = 3, 4, 5, ... that commutes with SU(5).
- $ightharpoonup Z_N^R [SU(3)_C]^2$  and  $Z_N^R [SU(2)_L]^2$  anomaly coefficients:  $\mathcal{A}_R^{(C)} \stackrel{(N)}{=} \mathcal{A}_R^{(L)} \stackrel{(N)}{=} -6$ .
- ▶ Given solely the MSSM particle content, only  $Z_3^R$  and  $Z_6^R$  anomaly-free.

## An Anomaly-Free Discrete R Symmetry for the MSSM (II)



 $N \neq 3,6 \Rightarrow$  Extra matter sector required. *Natural* consequende of gauged  $Z_N^R$ .

- ▶ Introduce k pairs of vector-like quark & anti-quark fields,  $Q_i \sim \mathbf{5}_i$  &  $\bar{Q}_i \sim \mathbf{5}_i^*$ .
- R charges such that  $k(r_Q + r_{\bar{Q}} 2) \stackrel{(N)}{=} +6$ . In most cases,  $r_Q + r_{\bar{Q}} \neq 0, 2$ .
- ▶  $W_Q^{\text{ren}} = 0$ .  $\Rightarrow$  global  $SU(k)_Q^V \times SU(k)_Q^A \times U(1)_Q^V \times U(1)_Q^A$  flavour symmetry.

## Rendering the Extra Quark Flavours Massive

Couple new matter sector to SM singlet *P* that acquires VEV above the EW scale:

$$W_Q \supset \frac{\lambda_i}{M_{\rm Pl}^{n-1}} P^n \left( Q \bar{Q} \right)_i, \quad m_{Q_i} = \frac{\lambda_i}{M_{\rm Pl}^{n-1}} \langle P \rangle^n, \quad n = 1, 2.$$

- $\triangleright$  *nk* possible values for  $r_P$ , the *R* charge of *P*, for each combination of *N*, *n*, *k*.
- Add singlets  $\bar{P}$  and X with  $r_{\bar{P}} = -r_P$  and  $r_X = 2$ . Restrict to values of  $r_P$  s. t.

$$W_P^{\rm ren} = \kappa X \left[ \frac{\Lambda^2}{2} - P \bar{P} \right], \quad \langle X \rangle \sim m_{3/2}, \quad \langle P \rangle = \frac{\Lambda}{\sqrt{2}} e^{A/\Lambda}, \quad \langle \bar{P} \rangle = \frac{\Lambda}{\sqrt{2}} e^{-A/\Lambda}.$$

▶ Global  $U(1)_P$ . Coupling b/t new matter and new singlet sector  $\rightarrow U(1)_{PQ}$ ,

$$U(1)_P \times U(1)_Q^V \times U(1)_Q^A \to U(1)_{PQ} \times U(1)_Q^V, \quad q_P = 1, \quad q_{\bar{P}} = -1.$$

- ▶ Colour anomaly:  $A_{PQ} = k q_{O\bar{O}} = k (-n)$ . Remniscent of KSVZ axion model.
- $ightharpoonup q_Q$  and  $q_{\bar{Q}}$  eventually fixed by coupling to MSSM (e.g.  $\bar{Q}\mathbf{10}H_d$  or  $\bar{P}\bar{Q}\mathbf{10}H_d$ ).

#### Generation of the MSSM $\mu$ Term

 $W_{\mu} = \mu H_u H_d$  forbidden by  $Z_N^R$ . Generated during / after R symmetry breaking.

- ▶ N = 4:  $K \supset g H_u H_d$ .  $\Rightarrow R$  breaking  $\rightarrow W \supset \frac{g}{M_{\rm Pl}^2} \langle W \rangle H_u H_d = g \, m_{3/2} H_u H_d$ .
- ▶  $N \neq 4$ : Couple standard model singlet S with  $r_s = -2$  to  $H_uH_d$ .

$$W_{S}^{\mathrm{ren}} = g_{H} H_{u} H_{d} S + m_{3/2}^{2} S + g_{X} m_{3/2} X S + g_{X^{2}} X^{2} S \quad \left( + m_{S} S^{2} \right) \, \left( + \lambda_{S} S^{3} \right) \, .$$

▶ In the PQ-breaking vacuum:  $\langle S \rangle = \mu/g_H \sim m_{3/2}$ .

#### Same low-energy phenomenology as the PQ-NMSSM and the nMSSM:

[Jeong, Shoji & Yamaguchi '12] [Panagiotakopoulos & Tamvakis '99; Panagiotakopoulos & Pilaftsis '01]

- ▶ Singlino  $\tilde{S}$  receives mass only from mixing with  $\tilde{H}_{u.d}^0$ .  $\Rightarrow$  Lightest neutralino.
- ▶ Contributions to  $m_{h^0}$  of a few GeV from singlino loops, if  $\tilde{H}^0_{u,d}$  are light.
- ▶ BR  $(h^0 \to \tilde{S}\tilde{S}) > 0.5$  at small tan  $\beta$ .  $\Rightarrow$  Soon tested at LHC-13 / LHC-14.

Our model: MSSM + extra **5**'s and **5**\*'s. + singlets P,  $\bar{P}$ , X + singlet S.

Minimal Extension of the MSSM with a Peccei-Quinn Symmetry

2 Phenomenological Constraints

#### Bounds on the Number of Extra Matter Multiplets

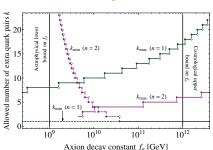
Require unification of the SM gauge couplings at the perturbative level,

$$g_{\mathrm{GUT}}(m_{Q_i}, k) \leq \sqrt{4\pi} \ \Rightarrow \ k_{\mathrm{max}} = k_{\mathrm{max}} \left( f_a, n \right),$$

and consistency with direct searches for heavy vector-like down-type quarks, [ATLAS, 14.3fb<sup>-1</sup> at √s = 8 TeV, assuming a dominant coupling to the third generation of SM quarks via the operator \(\tilde{O}10H\_{rd}\)]

$$m_{Q_i} \propto |\mathcal{A}_{PQ}|^n \propto k^n$$
,  $m_{Q_i} \geq M_Q^{\min} = 590 \,\text{GeV} \Rightarrow k_{\min} = k_{\min} (f_a, n)$ .

#### Solve RGEs including the new matter:



 $k_{\min}$  and  $k_{\max}$  translate into lower bounds on the axion decay constant  $f_a$ :

$$g_{\mathrm{GUT}}(f_a^{\mathrm{min,p}},n,k) = \sqrt{4\pi}, \ M_Q(f_a^{\mathrm{min,m}},n,k) = M_Q^{\mathrm{min}}.$$

$$\max\left\{f_a^{\min,p},f_a^{\min,m}\right\}\leq f_a.$$

$$f_a^{\text{min},i} = f_a^{\text{min},i}(k,n)\,, \quad i = p,m\,.$$

## Shifts in the QCD Vacuum Angle

Higher-dim. operators explicitly break the  $U(1)_{PQ}$ , Most relevant operators in W:

$$W \supset \frac{P^{p}S^{s}}{p!\,s!\,M_{Pl}^{c}},\,\, \frac{\bar{P}^{\bar{p}}S^{s}}{\bar{p}!\,s!\,M_{Pl}^{c}},\,\, \frac{m_{3/2}^{m}P^{p}X^{x}}{p!\,x!\,M_{Pl}^{c}},\,\, \frac{m_{3/2}^{m}\bar{P}^{\bar{p}}X^{x}}{\bar{p}!\,x!\,M_{Pl}^{c}}\,,\,\,\, r_{P}(p-\bar{p}) + 2(m+x-s) \stackrel{\text{\tiny (N)}}{=} 2\,.$$

Non-standard contributions to the axion potential (from *F*- and *A*-terms):

$$\Delta V_a = M^4 \cos \left( p \frac{a}{\sqrt{2}\Lambda} \right), \quad M = M \left( N, n, k, f_a, m_{3/2}, \langle S \rangle, \langle X \rangle \right)$$

These distortions of  $V_a$  induce shifts in the axion VEV,  $\langle a \rangle = \left( \bar{\theta} + \Delta \bar{\theta} \right) f_a$ :

$$\Delta \bar{\theta} \sim \frac{\rho}{|\mathcal{A}_{\rm PQ}|} \frac{M^4}{\Lambda_{\rm QCD}^4} \leq 10^{-10} \ \Rightarrow \ M^4 \leq 10^{-10} \frac{|\mathcal{A}_{\rm PQ}|}{\rho} \Lambda_{\rm QCD}^4 \ \Rightarrow \ \mathit{f}_a \leq \mathit{f}_a^{max} \,.$$

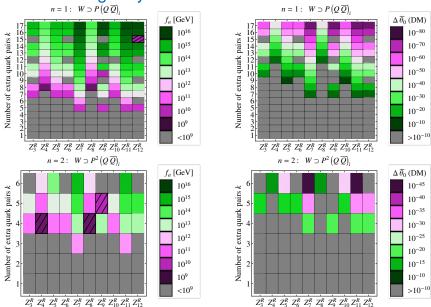
$$f_a \leq \min\left\{f_a^{\max,S}, f_a^{\min,X}\right\}.$$

$$10^9 \text{GeV} \lesssim f_a \lesssim 10^{12} \text{GeV}$$
.

 $(N, n, k, r_P)$  viable if window of viable  $f_a$ .

▶ We scan 1950 combinations of *N*, *n*, *k* and  $r_P$  for  $m_{3/2} = \langle S \rangle = \langle X \rangle = 1 \text{ TeV}$ .

## Phenomenologically Viable Scenarios



Minimal Extension of the MSSM with a Peccei-Quinn Symmetry

Phenomenological Constraints

## The PQ Symmetry from a Gauged Discrete R Symmetry

Problem: Anomalous global  $U(1)_{PQ}$ , required for the axion solution of the strong CP problem, expected to be broken by quantum gravity effects.

Idea: Approximate accidental  $U(1)_{PQ}$  due to exact gauged discrete R symmetry.

- New matter sector in order to render the  $Z_N^R$  anomaly-free,  $Q_i \& \bar{Q}_i$ .
- New singlet sector in order to provide masses to the new matter, P,  $\bar{P}$  & X.
- ▶ Singlet *S* to generate the MSSM  $\mu$  term.

#### Phenomenological constraints on N, n, k, $r_P$ , $f_a$ based on:

- Lower bound on the mass of heavy down-type quarks,  $M_Q^{\min} = 590 \, \mathrm{GeV}$ .
- SM gauge coupling unification at the perturbative level,  $g_{GUT} \leq \sqrt{4\pi}$ .
- ▶ Not too large a shift in the QCD vacuum angle,  $\bar{\theta} < 10^{-10}$ .
- ►  $f_a$  within astrophysically viable window,  $10^9 \text{GeV} \lesssim f_a \lesssim 10^{12} \, \text{GeV}$ .

Result: Large landscape of viable solutions. Lower bounds on  $\bar{\theta}$  in case of axion DM.

## The PQ Symmetry from a Gauged Discrete R Symmetry

Problem: Anomalous global  $U(1)_{PQ}$ , required for the axion solution of the strong CP problem, expected to be broken by quantum gravity effects.

Idea: Approximate accidental  $U(1)_{PQ}$  due to exact gauged discrete R symmetry.

- New matter sector in order to render the  $Z_N^R$  anomaly-free,  $Q_i \& \bar{Q}_i$ .
- New singlet sector in order to provide masses to the new matter, P,  $\bar{P}$  & X.
- ▶ Singlet S to generate the MSSM  $\mu$  term.

#### Phenomenological constraints on N, n, k, $r_P$ , $f_a$ based on:

- ▶ Lower bound on the mass of heavy down-type quarks,  $M_Q^{\min} = 590 \, \mathrm{GeV}$ .
- SM gauge coupling unification at the perturbative level,  $g_{GUT} \le \sqrt{4\pi}$ .
- Not too large a shift in the QCD vacuum angle,  $\bar{\theta} < 10^{-10}$ .
- ▶  $f_a$  within astrophysically viable window,  $10^9 \text{GeV} \lesssim f_a \lesssim 10^{12} \text{GeV}$ .

Result: Large landscape of viable solutions. Lower bounds on  $\bar{\theta}$  in case of axion DM.

#### Thank you for your attention!