

Electroweak precision fit and model-independent constraints on new physics

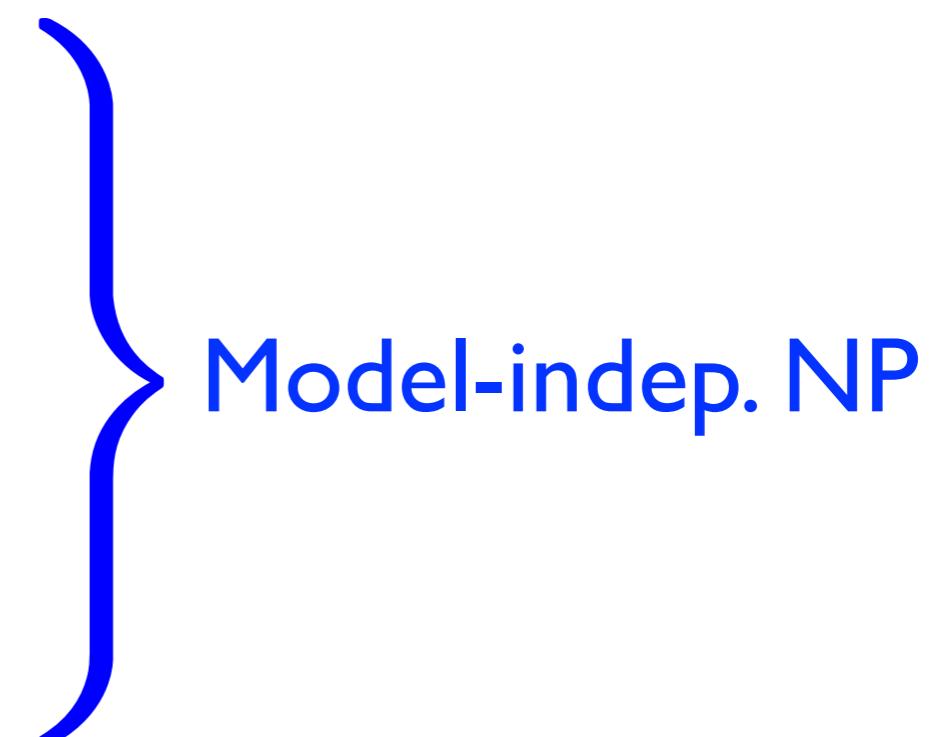
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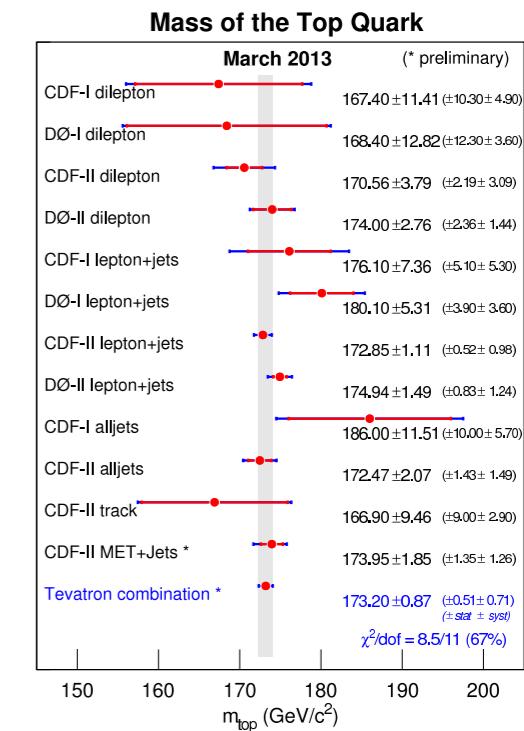
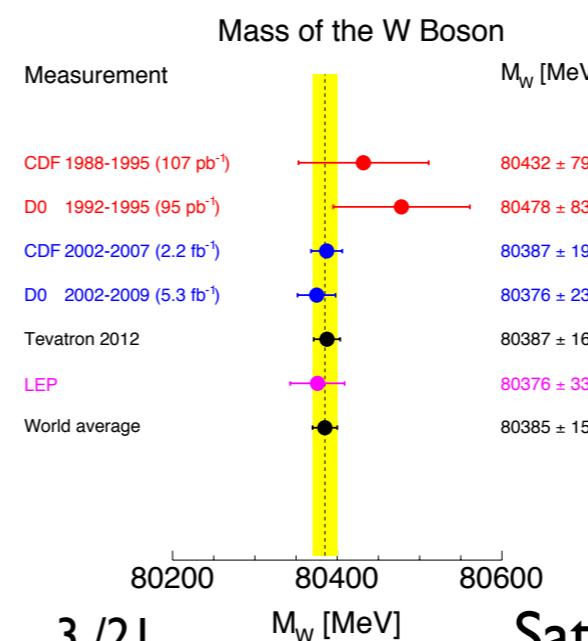
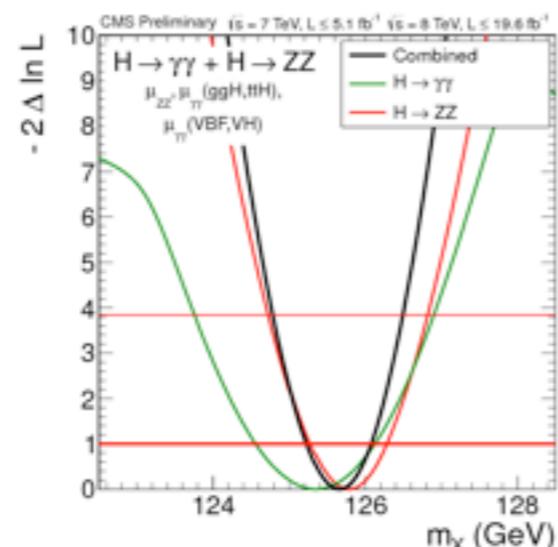
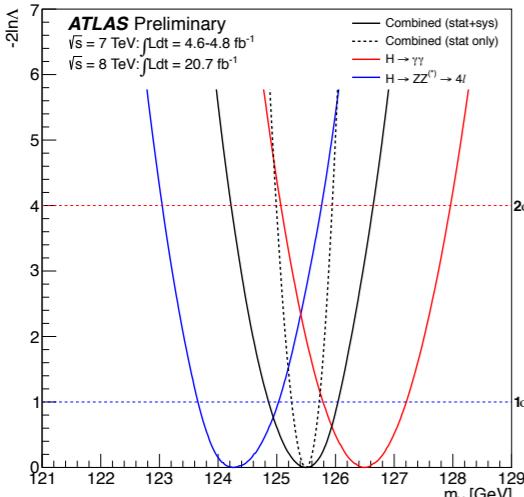
*with Marco Ciuchini, Enrico Franco and Luca Silvestrini,
JHEP 08 (2013) 106 [arXiv:1306.4644[hep-ph]]*

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- Model-indep. NP

1. Introduction

- Electroweak precision fits have played a key role in constraining NP.
- The precise measurement of the Higgs mass at LHC as well as those of the W and top masses at Tevatron make improvement in EW fits.
- It is therefore phenomenologically relevant to reassess the constraining power of EW fits in the light of the recent exp. and theo. improvements.



EW precision observables

$$\begin{aligned}\mathcal{L} &= \frac{e}{2s_W c_W} Z_\mu \sum_f \bar{f} \left(\textcolor{red}{g_V^f} \gamma_\mu - \textcolor{red}{g_A^f} \gamma_\mu \gamma_5 \right) f, \\ &= \frac{e}{2s_W c_W} Z_\mu \sum_f \bar{f} \left[\textcolor{red}{g_L^f} \gamma_\mu (1 - \gamma_5) + \textcolor{red}{g_R^f} \gamma_\mu (1 + \gamma_5) \right] f, \\ &= \frac{e}{2s_W c_W} \sqrt{\rho_Z^f} Z_\mu \sum_f \bar{f} \left[(I_3^f - 2Q_f \kappa_Z^f s_W^2) \gamma^\mu - I_3^f \gamma^\mu \gamma_5 \right] f\end{aligned}$$

$$\rho_Z^f = \left(\frac{g_A^f}{I_3^f} \right)^2 \quad (= 1 \text{ at tree level})$$

$$\kappa_Z^f = \frac{1}{4|Q_f|s_W^2} \left(1 - \frac{g_V^f}{g_A^f} \right)$$

$$A_{\text{LR}}^{0,f} = \mathcal{A}_f = \frac{2 \operatorname{Re} \left(g_V^f / g_A^f \right)}{1 + \left[\operatorname{Re} \left(g_V^f / g_A^f \right) \right]^2} \quad A_{\text{FB}}^{0,f} = \frac{3}{4} \mathcal{A}_e \mathcal{A}_f \quad (f = \ell, c, b)$$

$$P_{\tau}^{\text{pol}} = \mathcal{A}_{\tau} \qquad \qquad \sin^2 \theta_{\text{eff}}^{\text{lept}} = \text{Re}(\kappa_Z^\ell) s_W^2$$

$$\Gamma_f = \Gamma(Z \rightarrow f\bar{f}) \propto |\rho_Z^f| \left[\left| \frac{g_V^f}{g_A^f} \right|^2 R_V^f + R_A^f \right] \quad \rightarrow \quad \Gamma_Z, \quad \sigma_h^0 = \frac{12\pi}{M_Z^2} \frac{\Gamma_e \Gamma_h}{\Gamma_Z^2}, \quad R_\ell^0 = \frac{\Gamma_h}{\Gamma_\ell}, \quad R_{c,b}^0 = \frac{\Gamma_{c,b}}{\Gamma_h}$$

Theoretical status

- M_W has been calculated with **full EW two-loop** and leading higher-order contributions.

Awramik, Czakon, Freitas & Weiglein (04)

- $\sin^2 \theta_{\text{eff}}^f$ (equivalent to κ_Z^f) have been calculated with **full EW two-loop** (bosonic is missing for b) and leading higher-order contributions.

Awramik, Czakon & Freitas (06); Awramik, Czakon, Freitas & Kniehl (09)

- EW two-loop corrections to ρ_Z^f (normalization of Γ_f) were calculated in the **large mt expansion**.

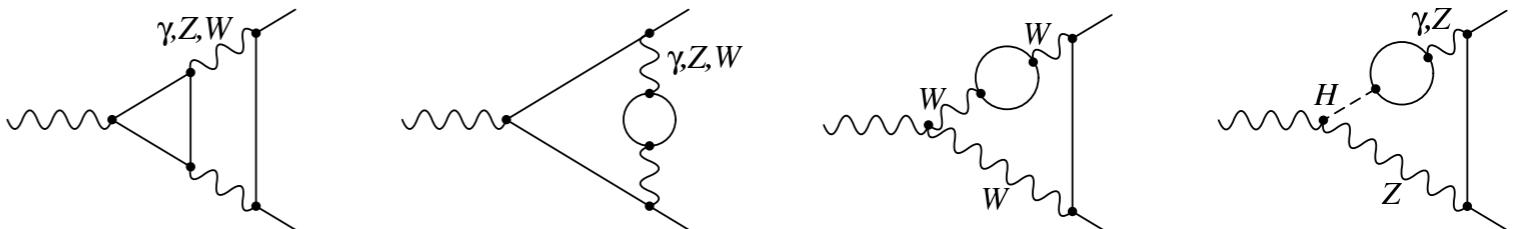
Barbieri et al. (92,93); Fleischer et al. (93,95); Degrassi et al. (96,97,00)

see also Sirlin; Marciano&Sirlin; Bardin et al; Djouadi&Verzegnassi; Djouadi; Kniehl; Halzen&Kniehl; Kniehl&Sirlin; Djouadi&Gambino; Avdeev et al; Chetyrkin et al; Freitas et al; Awramik&Czakon; Onishchenko&Veretin; Van der Bij et al; Faisst et al; Awramik et al

Theoretical status

- Recently, the **full fermionic EW two-loop corrections** to $R_b^0 = \Gamma_b/\Gamma_h$ beyond the large mt expansion have been calculated.

Freitas & Huang (12)



- Subleading power corrections are much larger than expected:
cf. $(\Delta R_b^0)_{\text{exp}} = \pm 0.00066$

$$R_b^0 = 0.21576 \rightarrow 0.21493 \quad (\Delta R_b^0 = -0.00083)$$

- The results are available only for $\Gamma_{u,d}/\Gamma_b$, which allow us to compute $R_c^0 = \Gamma_c/\Gamma_h$ in addition to R_b^0 .
- Subleading corr's: $\delta\rho_Z^{u(d)} - \delta\rho_Z^b \approx 0.0048(0.0044)$

$$\Gamma_f \propto |\rho_Z^f| \quad (\delta\Gamma_Z, \delta\sigma_h^0, \delta R_\ell^0)_{\text{exp}} \sim 0.1\%$$

→ We expect significant corrections also in Γ_Z , σ_h^0 , R_ℓ^0 !

2. SM fit

- Two options:
 1. “old R_b”: use ρ_Z^f in the large mt expansion
 2. “new R_b”: use $R_{c,b}^0$ from Freitas and Huang, introducing $\delta\rho_Z^{\nu,\ell,b}$ to account for possibly large unknown corrections.
- We adopt the **on-shell** renormalization scheme, and include up-to-date SM radiative corrections.
- We have developed **our own codes for EWPO**, tested against ZFITTER.
- We perform a **Bayesian** analysis with MCMC by using the Bayesian Analysis Toolkit (BAT) library.

cf. *Erler with GAPP for PDG* $\overline{\text{MS}}$, frequentist

*LEP EWWG with ZFITTER; Gfitter (Baak et al.);
Eberhardt et al. with ZFITTER; and others.....*

Caldwell, Kollar & Kroninger

on-shell, frequentist

Satoshi Mishima (Univ. of Rome)

SM input parameters

$$G_\mu = 1.1663787 \times 10^{-5} \text{ GeV}^{-2} \quad PDG$$

$$\alpha = 1/137.035999074 \quad PDG$$

$$\alpha_s(M_Z^2) = 0.1184 \pm 0.0006 \quad PDG \text{ excl. EW}$$

$$\Delta\alpha_{\text{had}}^{(5)}(M_Z^2) = 0.02750 \pm 0.00033$$

Burkhardt & Pietrzyk (II) (see also Davier et al(II);
Hagiwara et al(II) ;Jegerlehner(II))

measured with inclusive processes.

smaller uncertainty if using exclusive processes with pQCD, etc.

$$M_Z = 91.1875 \pm 0.0021 \text{ GeV} \quad LEP$$

$$m_t = 173.2 \pm 0.9 \text{ GeV} \quad \text{Tevatron (cf. LHC: } 173.3 \pm 1.4 \text{ GeV})$$

$$m_h = 125.6 \pm 0.3 \text{ GeV} \quad \text{ATLAS&CMS (naive average)}$$

+ $\delta\rho_Z^\nu$, $\delta\rho_Z^\ell$, $\delta\rho_Z^b$ for “new Rb”

SM fit with the old Rb

	Data	Fit	Indirect	Pull
$\alpha_s(M_Z^2)$	0.1184 ± 0.0006	0.1184 ± 0.0006	0.1193 ± 0.0027	+0.3
$\Delta\alpha_{\text{had}}^{(5)}(M_Z^2)$	0.02750 ± 0.00033	0.02740 ± 0.00026	0.02725 ± 0.00042	-0.5
M_Z [GeV]	91.1875 ± 0.0021	91.1878 ± 0.0021	91.197 ± 0.012	+0.8
m_t [GeV]	173.2 ± 0.9	173.5 ± 0.8	176.3 ± 2.5	+1.1
m_h [GeV]	125.6 ± 0.3	125.6 ± 0.3	97.3 ± 26.9	-0.9
M_W [GeV]	80.385 ± 0.015	80.367 ± 0.007	80.362 ± 0.007	-1.4
Γ_W [GeV]	2.085 ± 0.042	2.0891 ± 0.0006	2.0891 ± 0.0006	+0.1
Γ_Z [GeV]	2.4952 ± 0.0023	2.4953 ± 0.0004	2.4953 ± 0.0004	+0.0
σ_h^0 [nb]	41.540 ± 0.037	41.484 ± 0.004	41.484 ± 0.004	-1.5
$\sin^2 \theta_{\text{eff}}^{\text{lept}}(Q_{\text{FB}}^{\text{had}})$	0.2324 ± 0.0012	0.23145 ± 0.00009	0.23144 ± 0.00009	-0.8
P_τ^{pol}	0.1465 ± 0.0033	0.1476 ± 0.0007	0.1477 ± 0.0007	+0.3
\mathcal{A}_ℓ (SLD)	0.1513 ± 0.0021	0.1476 ± 0.0007	0.1471 ± 0.0008	-1.9
\mathcal{A}_c	0.670 ± 0.027	0.6682 ± 0.0003	0.6682 ± 0.0003	-0.1
\mathcal{A}_b	0.923 ± 0.020	0.93466 ± 0.00006	0.93466 ± 0.00006	+0.6
$A_{\text{FB}}^{0,\ell}$	0.0171 ± 0.0010	0.0163 ± 0.0002	0.0163 ± 0.0002	-0.8
$A_{\text{FB}}^{0,c}$	0.0707 ± 0.0035	0.0740 ± 0.0004	0.0740 ± 0.0004	+0.9
$A_{\text{FB}}^{0,b}$	0.0992 ± 0.0016	0.1035 ± 0.0005	0.1039 ± 0.0005	+2.8
R_ℓ^0	20.767 ± 0.025	20.735 ± 0.004	20.734 ± 0.004	-1.3
R_c^0	0.1721 ± 0.0030	0.17223 ± 0.00002	0.17223 ± 0.00002	+0.0
R_b^0	0.21629 ± 0.00066	0.21575 ± 0.00003	0.21575 ± 0.00003	-0.8

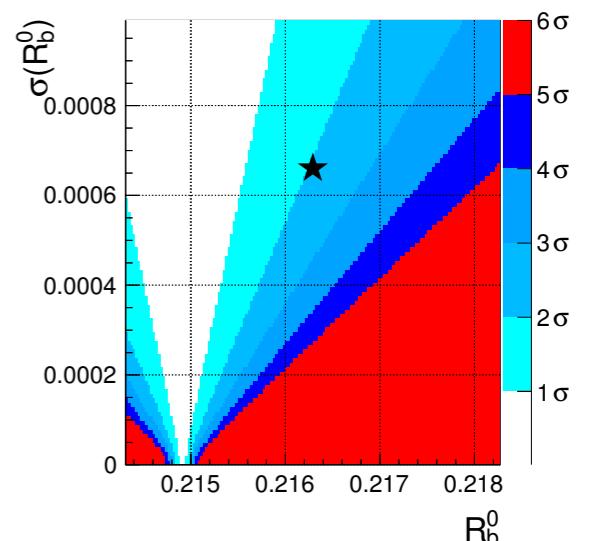
Fit: our fit results

Indirect: determined w/o using the corresponding experimental information

Pull: in units of standard deviations evaluated from the p.d.f.'s of "Data" and "Indirect"

SM fit with the new Rb

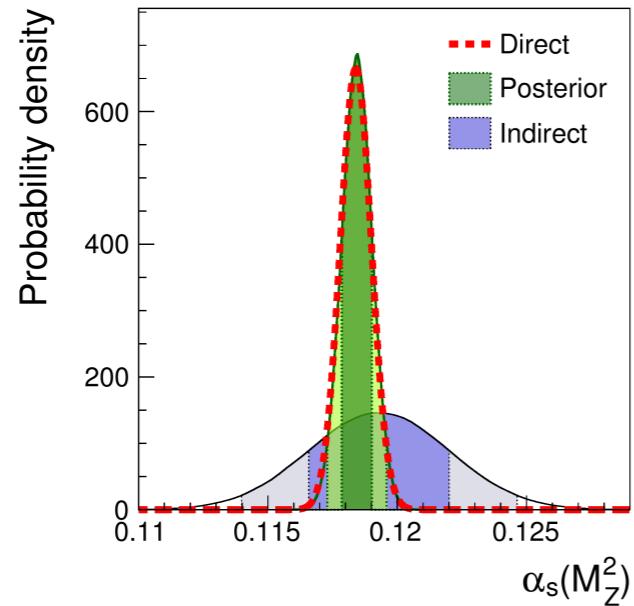
	Data	Fit	Indirect	Pull	
$\alpha_s(M_Z^2)$	0.1184 ± 0.0006	0.1184 ± 0.0006	0.078 ± 0.024	-1.9	← not precise!
$\Delta\alpha_{\text{had}}^{(5)}(M_Z^2)$	0.02750 ± 0.00033	0.02742 ± 0.00026	0.02728 ± 0.00043	-0.4	
M_Z [GeV]	91.1875 ± 0.0021	91.1878 ± 0.0021	91.204 ± 0.013	+1.2	
m_t [GeV]	173.2 ± 0.9	173.5 ± 0.8	175.7 ± 2.6	+0.9	
m_h [GeV]	125.6 ± 0.3	125.6 ± 0.3	98.5 ± 27.7	-0.8	
$\delta\rho_Z^\nu$	—	-0.0052 ± 0.0031	—	—	
$\delta\rho_Z^\ell$	—	-0.0002 ± 0.0010	—	—	
$\delta\rho_Z^b$	—	-0.0021 ± 0.0011	—	—	
M_W [GeV]	80.385 ± 0.015	80.366 ± 0.007	80.361 ± 0.007	-1.4	
Γ_W [GeV]	2.085 ± 0.042	2.0890 ± 0.0006	2.0890 ± 0.0006	+0.1	
Γ_Z [GeV]	2.4952 ± 0.0023	2.4952 ± 0.0023	—	—	
σ_h^0 [nb]	41.540 ± 0.037	41.539 ± 0.037	—	—	
$\sin^2\theta_{\text{eff}}^{\text{lept}}(Q_{\text{FB}}^{\text{had}})$	0.2324 ± 0.0012	0.23145 ± 0.00009	0.23145 ± 0.00009	-0.8	
P_τ^{pol}	0.1465 ± 0.0033	0.1476 ± 0.0007	0.1476 ± 0.0007	+0.3	
\mathcal{A}_ℓ (SLD)	0.1513 ± 0.0021	0.1476 ± 0.0007	0.1470 ± 0.0008	-1.9	
\mathcal{A}_c	0.670 ± 0.027	0.6681 ± 0.0003	0.6681 ± 0.0003	-0.1	
\mathcal{A}_b	0.923 ± 0.020	0.93466 ± 0.00006	0.93466 ± 0.00006	+0.6	
$A_{\text{FB}}^{0,\ell}$	0.0171 ± 0.0010	0.0163 ± 0.0002	0.0163 ± 0.0002	-0.8	
$A_{\text{FB}}^{0,c}$	0.0707 ± 0.0035	0.0739 ± 0.0004	0.0740 ± 0.0004	+0.9	
$A_{\text{FB}}^{0,b}$	0.0992 ± 0.0016	0.1034 ± 0.0005	0.1038 ± 0.0005	+2.7	
R_ℓ^0	20.767 ± 0.025	20.768 ± 0.025	—	—	
R_c^0	0.1721 ± 0.0030	0.17247 ± 0.00002	0.17247 ± 0.00002	+0.1	
R_b^0	0.21629 ± 0.00066	0.21492 ± 0.00003	0.21492 ± 0.00003	-2.1	← new Rb



- New Rb increases the pull from -0.8σ to -2.1σ .
- $\delta\rho_Z^{\nu,\ell,b}$ are of $O(10^{-3})$, compatible to Freitas & Huang.

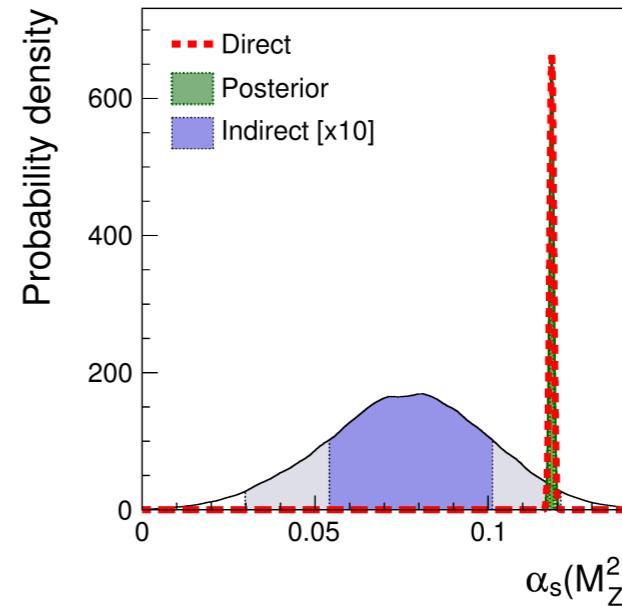
$\alpha_s(M_Z^2)$ from the EW fit

old Rb



$$0.1193 \pm 0.0027$$

new Rb

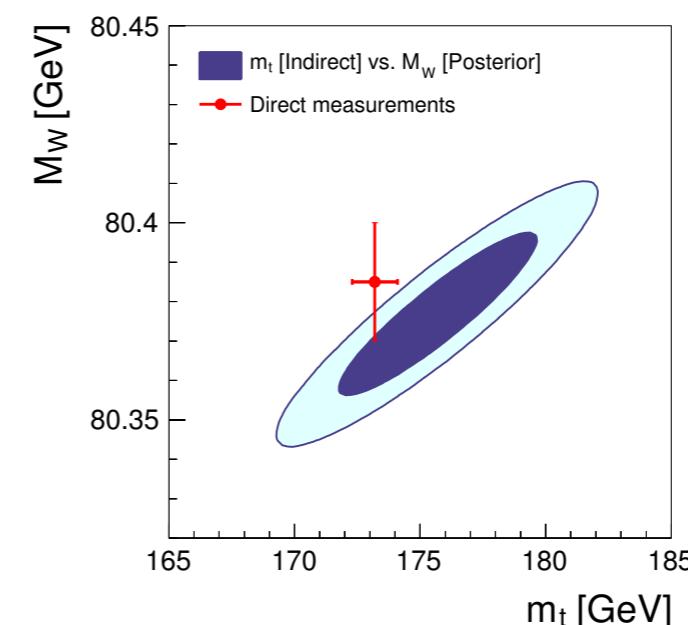
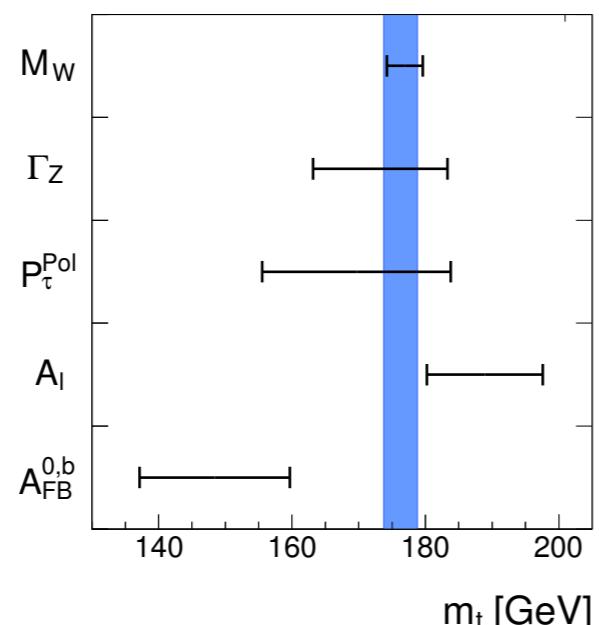
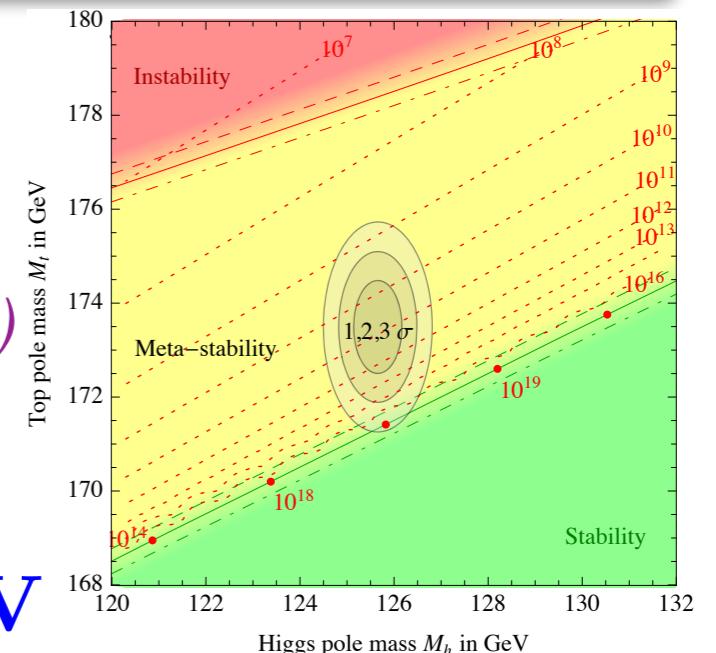


$$0.078 \pm 0.024$$

Not precise due to the inclusion of
the unknown subleading corrections!

Top mass vs. (meta-)stability

- The measurement of the top mass is crucial for testing the stability of the SM vacuum. *Degrassi et al.(12); Buttazzo et al.(13)*
 $m_t^{\text{pole}} < (171.36 \pm 0.46) \text{ GeV}$
- Tevatron pole(?) mass: **$173.2 \pm 0.9 \text{ GeV}$**
- Pole from MSbar: **$173.3 \pm 2.8 \text{ GeV}$**
- Indirect determination from EW fit: **$175.7 \pm 2.6 \text{ GeV}$**



On NP fits

“new Rb”

Significant subleading corrections in Γ_Z , σ_h^0 , R_ℓ^0



We do not use them in NP fits with the new $R_{c,b}^0$.

“old Rb”

For comparison, we also present fit results with the old $R_{c,b}^0$ and Γ_Z , σ_h^0 , R_ℓ^0 calculated using the large mt expansion.

3. Oblique parameters

- Suppose that dominant NP effects appear in the vacuum polarizations of the gauge bosons:

$$S = -16\pi \Pi'_{30}(0) = 16\pi \left[\Pi'^{\text{NP}}_{33}(0) - \Pi'^{\text{NP}}_{3Q}(0) \right]$$

$$T = \frac{4\pi}{s_W^2 c_W^2 M_Z^2} \left[\Pi'^{\text{NP}}_{11}(0) - \Pi'^{\text{NP}}_{33}(0) \right]$$

$$U = 16\pi \left[\Pi'^{\text{NP}}_{11}(0) - \Pi'^{\text{NP}}_{33}(0) \right]$$

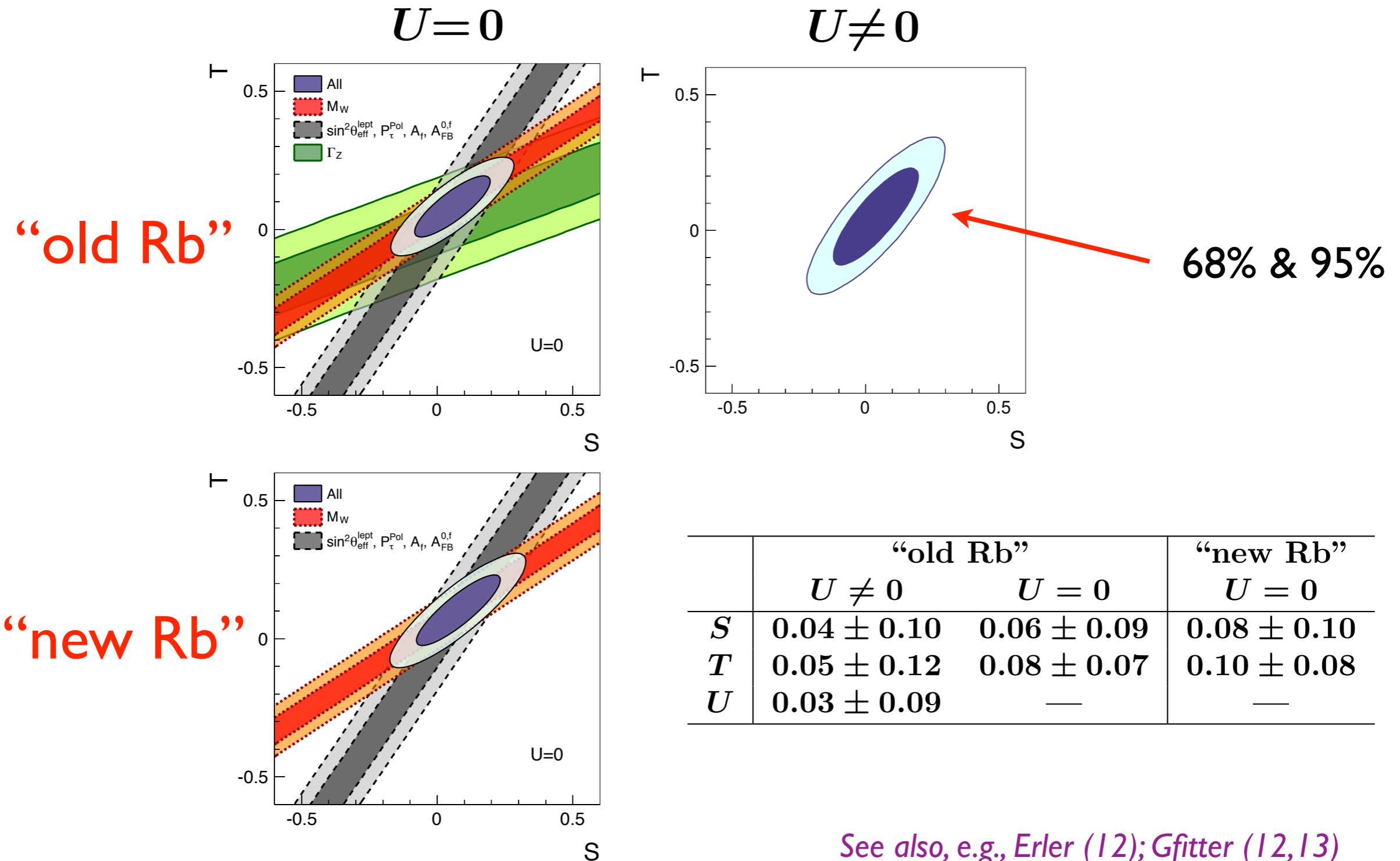
*Kennedy & Lynn (89);
Peskin & Takeuchi (90,92)*

- EWPO depend on the combinations:

$$\begin{aligned}\delta M_W, \delta \Gamma_W &\propto -\textcolor{red}{S} + 2c_W^2 \textcolor{red}{T} + \frac{(c_W^2 - s_W^2) \textcolor{red}{U}}{2s_W^2} \\ \delta \Gamma_Z &\propto -10(3 - 8s_W^2) \textcolor{red}{S} + (63 - 126s_W^2 - 40s_W^4) \textcolor{red}{T} \\ \text{others} &\propto \textcolor{red}{S} - 4c_W^2 s_W^2 \textcolor{red}{T}\end{aligned}$$

- Cannot use $\textcolor{red}{\Gamma}_Z$ with the new Rb.
 **cannot fit all S, T and U.**

Oblique parameters



4. HVV coupling

- only a Higgs below cutoff + custodial symmetry:

$$\mathcal{L} = \frac{v^2}{4} \text{Tr}(D_\mu \Sigma^\dagger D^\mu \Sigma) \left(1 + 2\textcolor{red}{a} \frac{h}{v} + \dots \right) + \dots$$

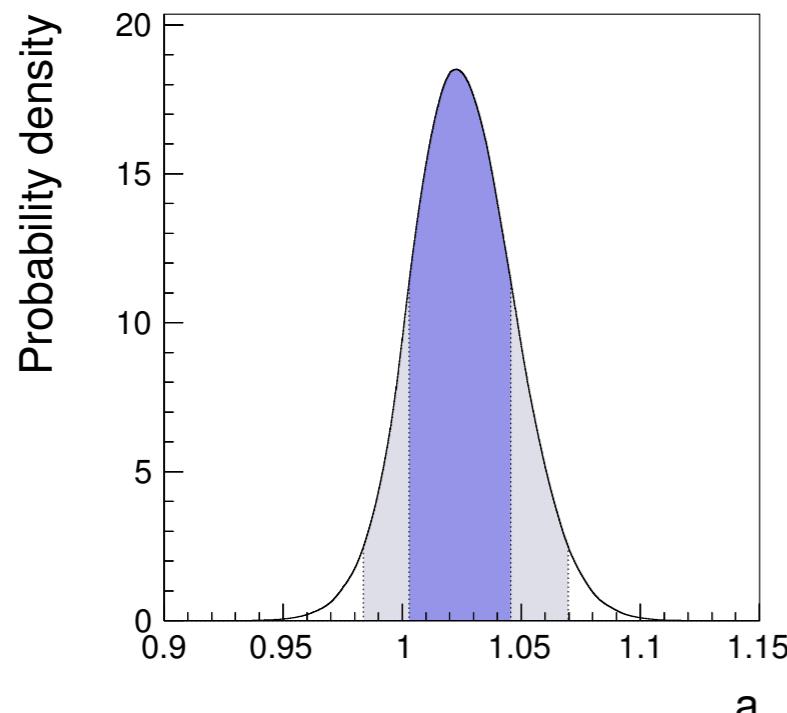
SM : $a = 1$

$$S = \frac{1}{12\pi} (1 - \textcolor{red}{a}^2) \ln \left(\frac{\Lambda^2}{m_h^2} \right)$$

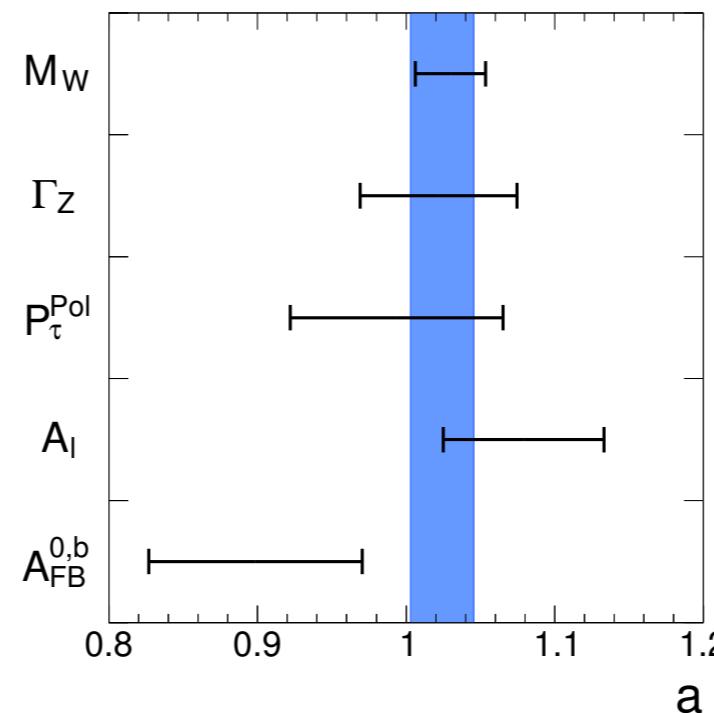
$$T = -\frac{3}{16\pi c_W^2} (1 - \textcolor{red}{a}^2) \ln \left(\frac{\Lambda^2}{m_h^2} \right)$$

$$\Lambda = 4\pi v / \sqrt{|1 - a^2|}$$

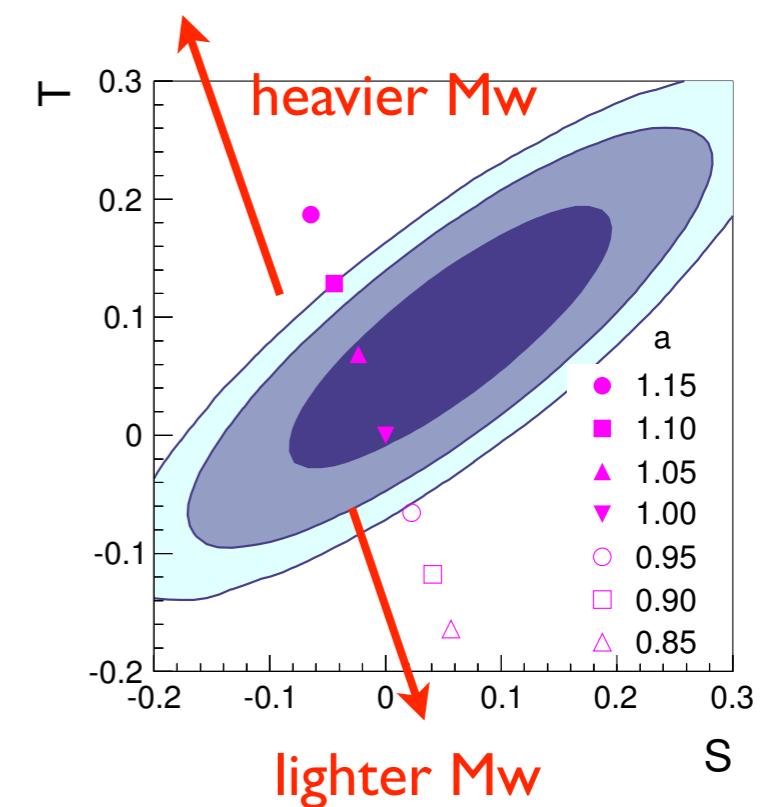
Barberi, Bellazzini, Rychkov & Varagnolo (07)



$$a = 1.024 \pm 0.021$$



“old Rb” (“new Rb” is similar)



Implication on composite Higgs models

- $a \in [0.98, 1.07] @ 95\%$
→ $\Lambda \gtrsim 16 \text{ TeV} @ 95\%$

- Composite Higgs models typically generate $a < 1$.

$$\xi = \left(\frac{v}{f}\right)^2 = 1 - a^2$$

$a > 1$ → $W_L W_L$ scattering is dominated by **isospin 2 channel**

Falkowski, Rychkov & Urbano (12)

- Extra light states are required to fix the EW fit.

Fermionic resonances can give desired shifts in S and T.

Grojean et al. (13); Azatov et al. (13)

- The result is less sensitive to the current LHC data.

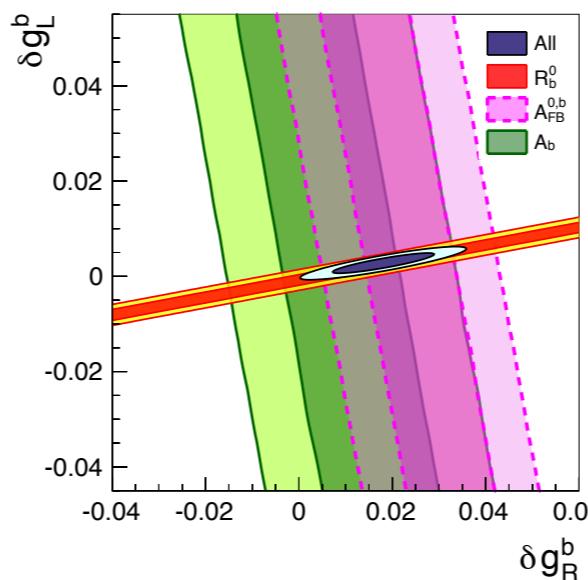
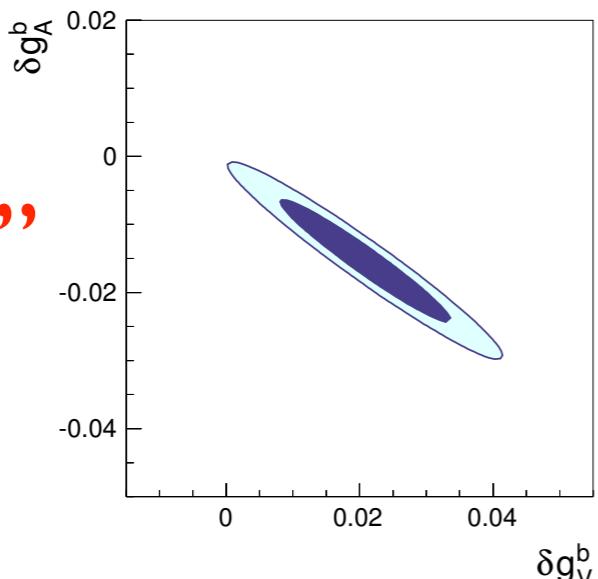
See, e.g., Azatov et al. (13); Falkowski et al. (13); Giardino et al. (13).

5. Zbb couplings

- Four solutions from Z-pole data, while two of them are disfavored by off Z-pole data for AFBb.
- The solution, closer to the SM:

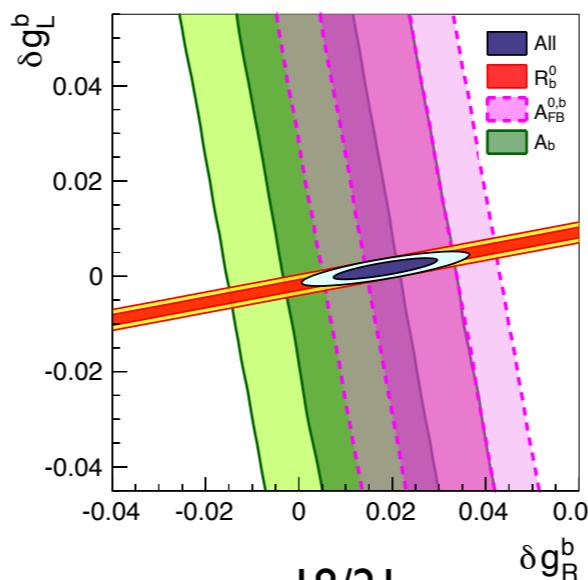
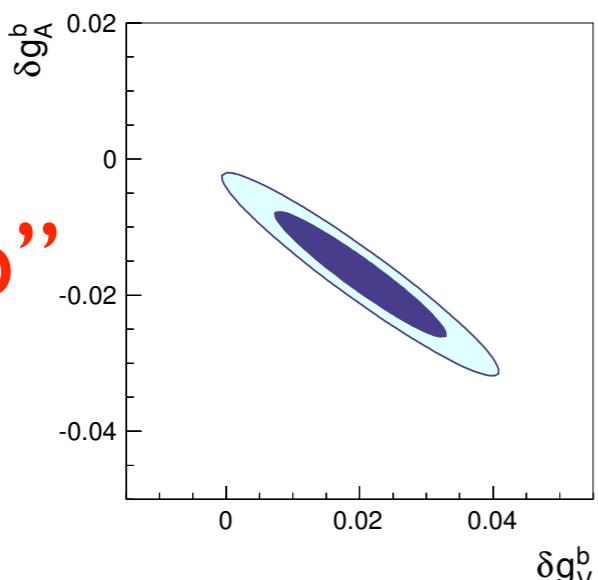
Choudhury et al. (02)

“old Rb”



	“old Rb”	“new Rb”
δg_R^b	0.018 ± 0.007	0.019 ± 0.007
δg_L^b	0.0028 ± 0.0014	0.0016 ± 0.0015
δg_V^b	0.021 ± 0.008	0.020 ± 0.008
δg_A^b	-0.015 ± 0.006	-0.017 ± 0.006

“new Rb”



See also Batell et al. (13)

6. Dim. 6 operators

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \sum_i \frac{C_i}{\Lambda^2} \mathcal{O}_i$$

Barbieri & Strumia (99)

$\mathcal{O}_{WB} = (H^\dagger \tau^a H) W_{\mu\nu}^a B^{\mu\nu},$	$\mathcal{O}_H = H^\dagger D_\mu H ^2,$
$\mathcal{O}_{LL} = \frac{1}{2} (\bar{L} \gamma_\mu \tau^a L)^2,$	$\mathcal{O}'_{HL} = i(H^\dagger D_\mu \tau^a H)(\bar{L} \gamma^\mu \tau^a L),$
$\mathcal{O}'_{HQ} = i(H^\dagger D_\mu \tau^a H)(\bar{Q} \gamma^\mu \tau^a Q),$	$\mathcal{O}_{HL} = i(H^\dagger D_\mu H)(\bar{L} \gamma^\mu L),$
$\mathcal{O}_{HQ} = i(H^\dagger D_\mu H)(\bar{Q} \gamma^\mu Q),$	$\mathcal{O}_{HE} = i(H^\dagger D_\mu H)(\bar{E} \gamma^\mu E),$
$\mathcal{O}_{HU} = i(H^\dagger D_\mu H)(\bar{U} \gamma^\mu U),$	$\mathcal{O}_{HD} = i(H^\dagger D_\mu H)(\bar{D} \gamma^\mu D).$



Switching on one operator at a time:

Coefficient	“old Rb”				“new Rb”			
	C_i/Λ^2 [TeV $^{-2}$] at 95%	$C_i = -1$	$C_i = 1$	Λ [TeV]	C_i/Λ^2 [TeV $^{-2}$] at 95%	$C_i = -1$	$C_i = 1$	Λ [TeV]
C_{WB}	[-0.0096, 0.0042]	10.2	15.4	10.2	[-0.0095, 0.0045]	10.3	15.0	10.2
C_H	[-0.030, 0.007]	5.8	12.1	5.8	[-0.031, 0.008]	5.7	11.5	5.8
C_{LL}	[-0.011, 0.019]	9.5	7.2	9.5	[-0.016, 0.023]	8.0	6.6	9.5
C'_{HL}	[-0.012, 0.005]	9.2	14.1	9.2	[-0.017, 0.009]	7.6	10.8	9.2
C'_{HQ}	[-0.010, 0.015]	10.2	8.2	10.2	[-0.40, 0.20]	1.6	2.2	10.2
C_{HL}	[-0.007, 0.010]	12.2	10.0	12.2	[-0.034, 0.022]	5.5	6.7	12.2
C_{HQ}	[-0.023, 0.046]	6.6	4.7	6.6	[-0.01, 0.11]	11.7	3.0	6.6
C_{HE}	[-0.014, 0.008]	8.4	11.1	8.4	[-0.029, 0.019]	5.9	7.2	8.4
C_{HU}	[-0.061, 0.087]	4.0	3.4	4.0	[-0.37, 0.08]	1.6	3.5	4.0
C_{HD}	[-0.15, 0.05]	2.6	4.6	2.6	[-1.1, -0.2]	1.0	—	2.6

Constraints on dim. 6 operators

Fit multiple op's simultaneously:

	“new Rb”
Coefficient	$C_i/\Lambda^2 \text{ [TeV}^{-2}\text{]} \text{ at 95\%}$
C_{WB}	[-0.009, 0.021]
C_H	[-0.068, 0.016]
C'_{HL}	[-0.029, 0.006]
C'_{HQ}	[-0.34, 0.31]
C_{HL}	—
C_{HQ}	[-0.07, 0.12]
C_{HE}	—
C_{HU}	[-0.26, 0.49]
C_{HD}	[-1.2, -0.2]
$C[\mathcal{A}_\ell]$	[-0.0021, 0.0050]
C'_{HL}	[-0.029, 0.006]
C_{HL}	—
C_{HE}	—
C_{HU_2}	[-0.32, 0.55]
C_{HD_3}	[-1.2, -0.2]
$C'_{HQ_1} + C'_{HQ_2}$	[-0.68, 0.61]
$C'_{HQ_2} - C_{HQ_2}$	[-0.67, 0.55]
$C'_{HQ_3} + C_{HQ_3}$	[-0.77, 0.63]
$C[\mathcal{A}_\ell]$	[-0.0021, 0.0050]
$C[\Gamma_{uds}]$	[-0.42, 0.43]

w/o quark-flavor universality:

Coefficient	“new Rb”	
	$C_i/\Lambda^2 \text{ [TeV}^{-2}\text{]} \text{ at 95\%}$	$\Lambda \text{ [TeV]}$
	$C_i = -1$	$C_i = 1$
C'_{HQ_1}	[-0.19, 0.01]	2.3
C'_{HQ_2}	[-0.20, 0.01]	2.3
C'_{HQ_3}, C_{HQ_3}	[0.00, 0.10]	15.6
C_{HQ_1}	[-1.9, 0.1]	0.7
C_{HQ_2}	[-0.25, 0.15]	2.0
C_{HU_1}	[-0.97, 0.03]	1.0
C_{HU_2}	[-0.39, 0.21]	1.6
C_{HD_1}, C_{HD_2}	[-0.1, 1.9]	3.8
C_{HD_3}	[-0.66, -0.13]	1.2

Recent exp. improvements strengthen the bounds on NP!

7. Summary

- We have updated the EW fit with the recent exp. and theo. improvements, discussing in detail the impact of the recently computed two-loop fermionic EW corrections to the $Z\bar{f}\bar{f}$ couplings.
- We have derived constraints on oblique parameters, HVV coupling, $Zb\bar{b}$ couplings and dim. 6 operators.
- The constraining power of the EW fit have been improved.
- Full two-loop calculations of Γ_Z , σ_h^0 , R_ℓ^0 need to be completed.
- Future data (LHC, ILC, LEP3/TLEP,...) strengthen greatly the power of the EW fit.

Backup

Comparison to ZFITTER

- For a given set of the input parameters,

	ZFITTER	OURS	$\frac{\text{OURS} - \text{ZFITTER}}{\text{ZFITTER}} * 100$	Exp uncertainty
M_W	80.362216	80.362499	0.00035 %	0.02 %
Γ_W	2.0906748	2.0887391	-0.093 %	2.0 %
Γ_Z	2.4953142	2.4951814	-0.0053 %	0.09 %
σ_h^0	41.479103	41.483516	0.011 %	0.09 %
$\sin^2 \theta_{\text{eff}}^{\text{lept}}(Q_{\text{FB}}^{\text{had}})$	0.23149326	0.23149297	-0.00012 %	0.52 %
P_{τ}^{Pol}	0.14724705	0.14724926	0.0015 %	2.2 %
A_{ℓ}	0.14724705	0.14724926	0.0015 %	1.4 %
A_c	0.66797088	0.66799358	0.0034 %	4.0 %
A_b	0.93460981	0.93464051	0.0033 %	2.2 %
$A_{\text{FB}}^{0,\ell}$	0.016261269	0.016261758	0.0030 %	5.5 %
$A_{\text{FB}}^{0,c}$	0.073767554	0.073771169	0.0049 %	5.0 %
$A_{\text{FB}}^{0,b}$	0.10321390	0.10321884	0.0048 %	1.6 %
R_{ℓ}^0	20.739702	20.735130	-0.022 %	0.12 %
R_c^0	0.17224054	0.17222362	-0.0098 %	1.7 %
R_b^0	0.21579927	0.21578277	-0.0077 %	0.31 %

Our results are in agreement with ZFITTER v6.43.

Impact of parametric uncertainties

	Prediction	α_s	$\Delta\alpha_{\text{had}}^{(5)}$	M_Z	m_t
M_W [GeV]	80.362 ± 0.008	± 0.000	± 0.006	± 0.003	± 0.005
Γ_Z [GeV]	2.4951 ± 0.0005	± 0.0003	± 0.0003	± 0.0002	± 0.0002
$P_{\tau}^{\text{pol}} = \mathcal{A}_\ell$	0.1472 ± 0.0009	± 0.0000	± 0.0009	± 0.0001	± 0.0002
$A_{\text{FB}}^{0,b}$	0.1032 ± 0.0007	± 0.0000	± 0.0006	± 0.0001	± 0.0002
R_b^0	0.21493 ± 0.00004	± 0.00001	± 0.00000	± 0.00000	± 0.00003

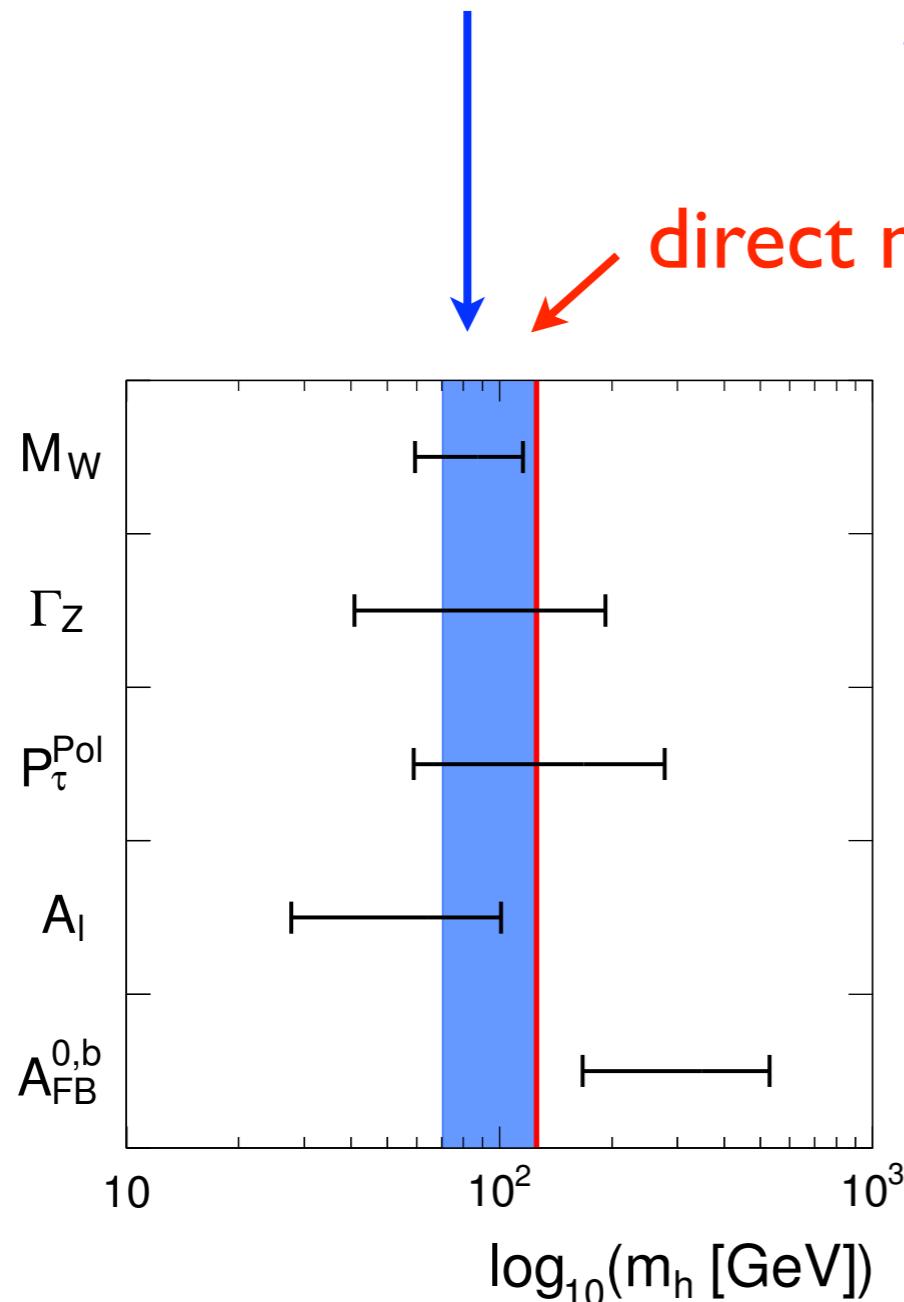
- $\Delta\alpha_{\text{had}}^{(5)}(M_Z^2)$ and m_t are the most important sources of parametric uncertainty.
- The theoretical uncertainty from missing higher-order corrections has been estimated as $\delta M_W^{\text{theo}} \sim 4$ MeV.

Awramik et al. (04)

Individual constraints on the Higgs mass

indirect determination from the EW fit:

$$m_h = 97.3 \pm 26.9 \text{ GeV}$$



direct measurement at LHC (ATLAS & CMS):

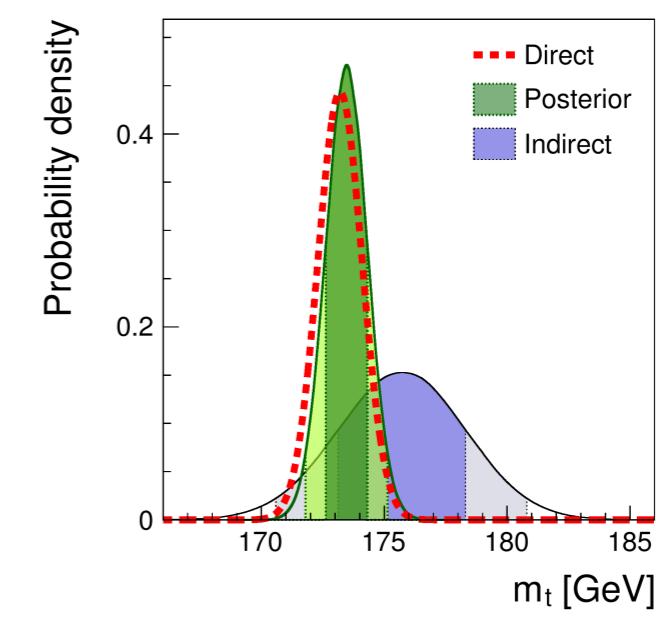
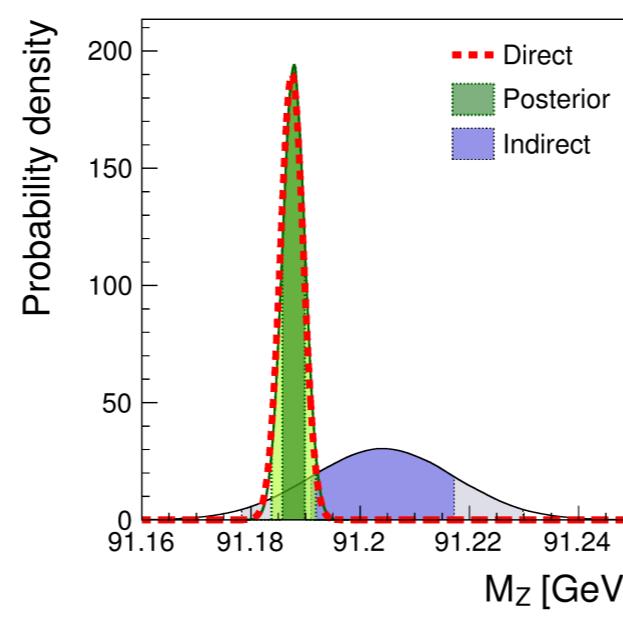
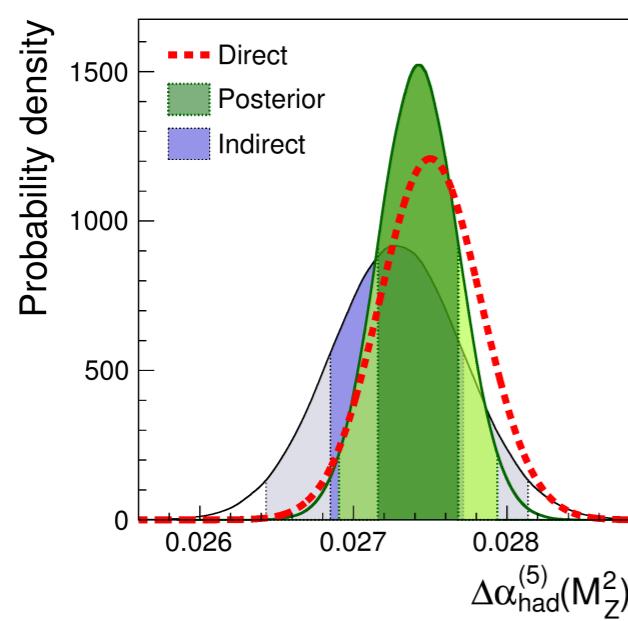
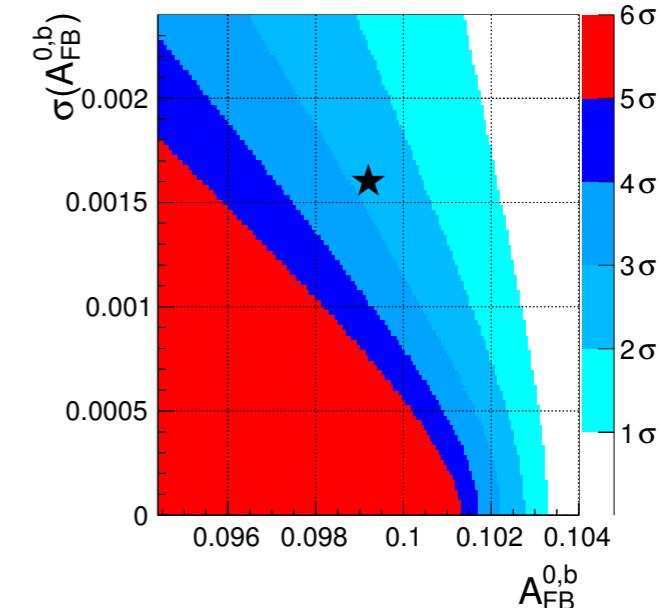
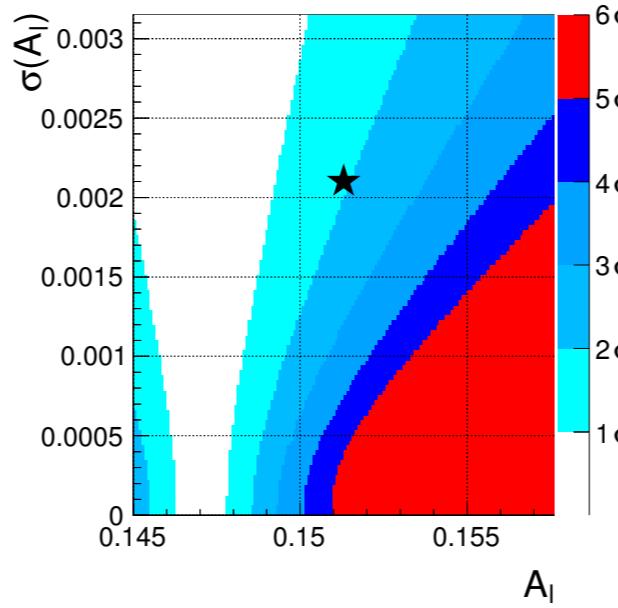
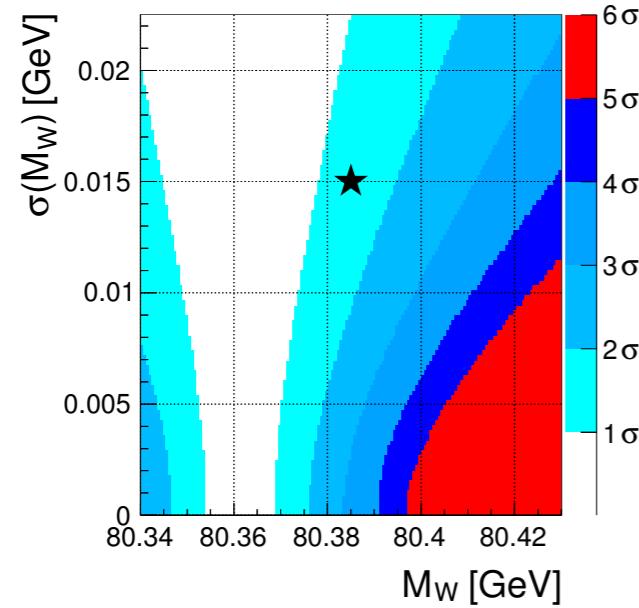
$$m_h = 125.6 \pm 0.3 \text{ GeV}$$

- M_W gives the most stringent constraint.

- Tension between $A_I(\text{SLD})$ and A_{FB}^b .

“old Rb”

Other plots in the SM fit



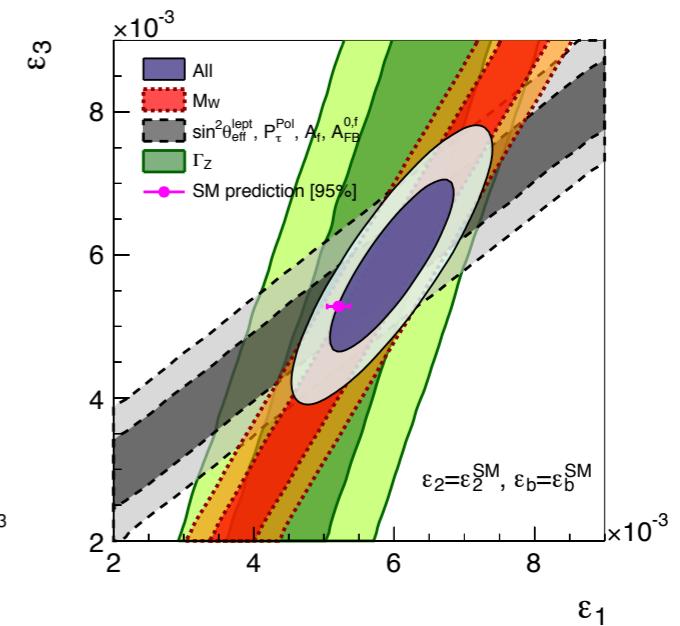
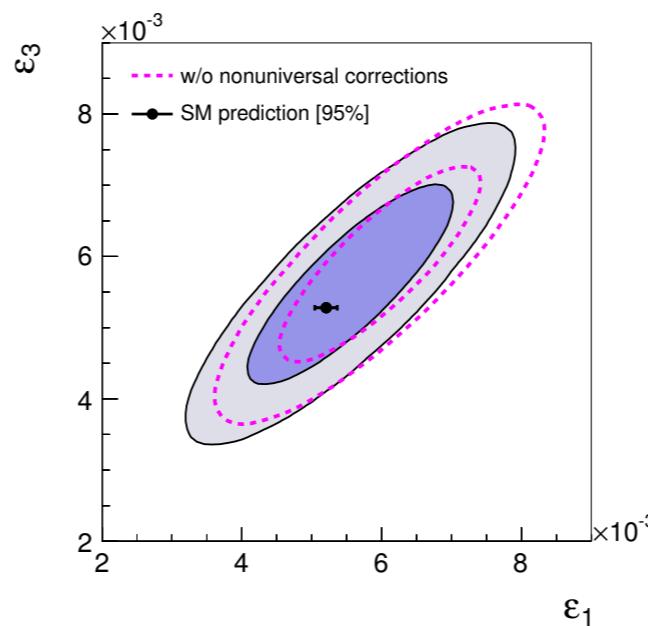
Constraints on the epsilon parameters

- Unlike STU, the epsilon parameters involve SM contributions.

Altarelli et al. (91,92,93)

Flavour non-universal VCs in the SM have to be taken into account.

Parameter	“old Rb”	
	$\epsilon_{1,2,3,b}$ fit	$\epsilon_{1,3}$ fit
$\epsilon_1 [10^{-3}]$	5.6 ± 1.0	6.0 ± 0.6
$\epsilon_2 [10^{-3}]$	-7.8 ± 0.9	—
$\epsilon_3 [10^{-3}]$	5.6 ± 0.9	5.9 ± 0.8
$\epsilon_b [10^{-3}]$	-5.8 ± 1.3	—



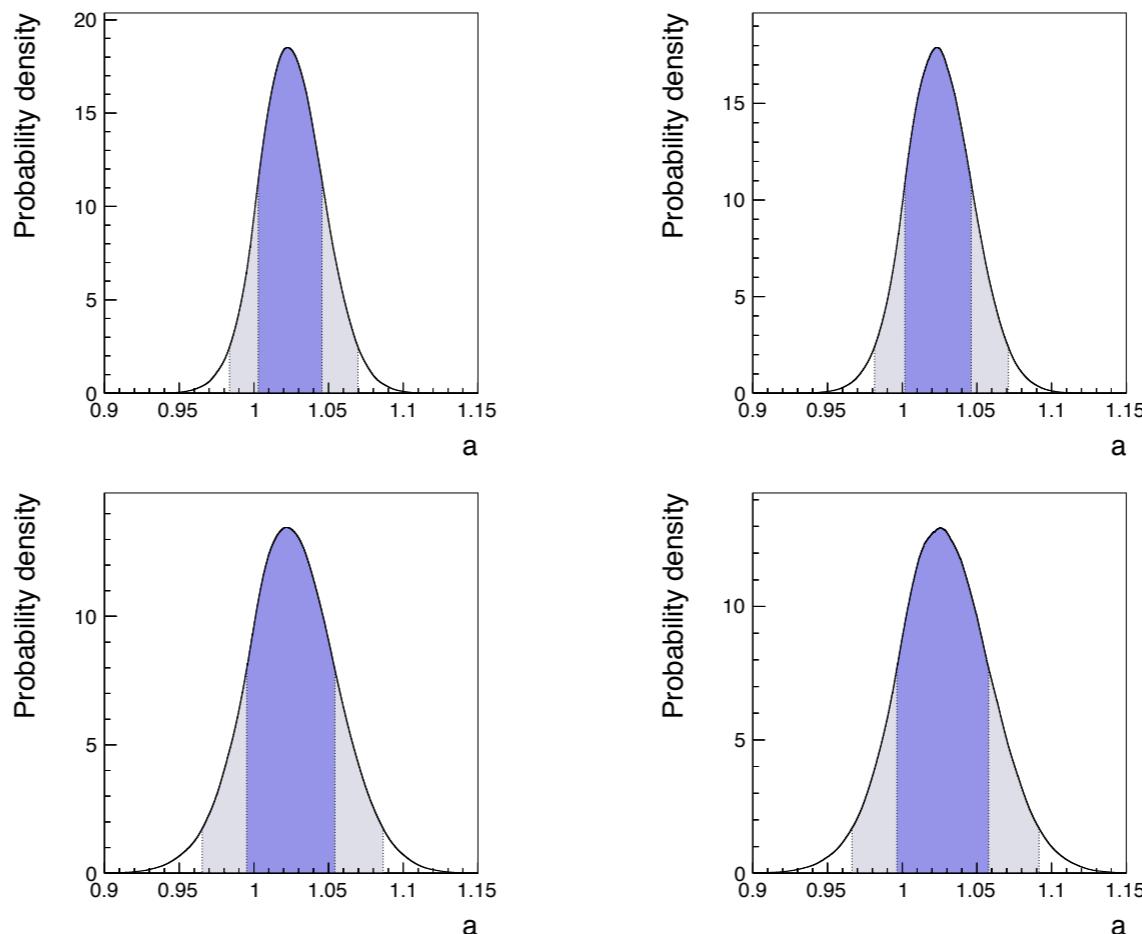
Top mass uncertainty for HVV

- If using the pole mass derived from the $\overline{\text{MS}}$ mass,

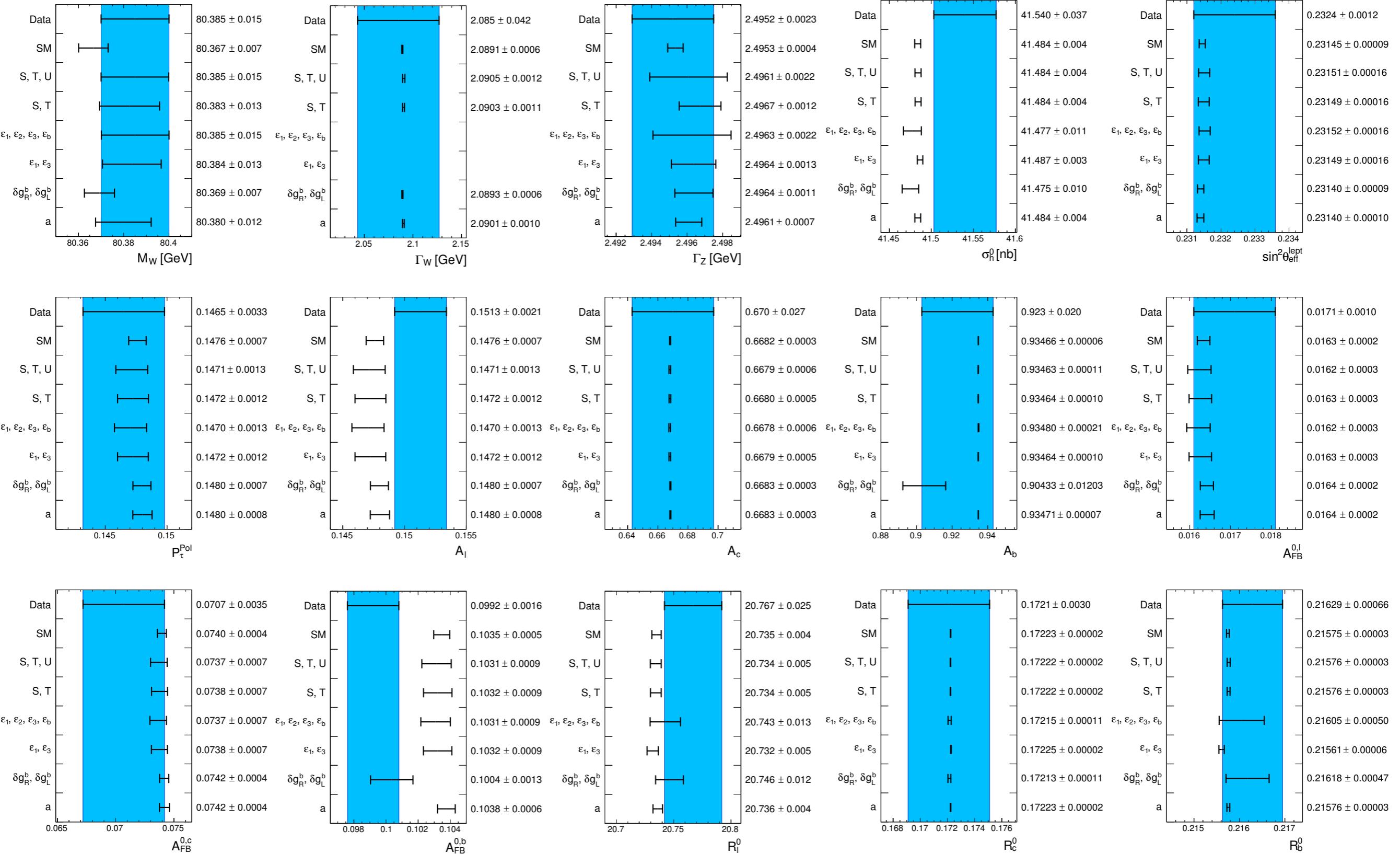
$$m_t^{\overline{\text{MS}}} \text{ from total } \sigma_{t\bar{t}} \rightarrow m_t^{\text{pole}} = 173.3 \pm 2.8 \text{ GeV}$$

m_t [GeV]	“old Rb”	“new Rb”
173.2 ± 0.9	1.024 ± 0.021	1.024 ± 0.022
173.3 ± 2.8	1.025 ± 0.030	1.027 ± 0.031

Alekhin et al. (12)



Summary of fit results w/o new Rb



Summary of fit results with new Rb

