Baryonic R-parity violation and Grand Unification

SUSY 2013, ICTP Trieste

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Based on arXiv:1305.7034 in collaboration with: Marco Nardecchia (CP3-Origins, Odense) and Andrea Romanino (SISSA, Trieste)



- R-parity violating SUSY @ LHC
- GUT and (baryonic) RPV: the problem
- A simple SO(10) model



MSSM & R-parity

• Most generic superpotential

$$W_{RPC} = \mu h_u h_d + y_{ij}^e e_i^c l_j h_d + y_{ij}^d q_i d_j^c h_d + y_{ij}^u q_j u_j^c h_u$$

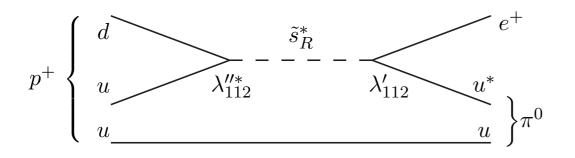


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$$W_{RPV} = \mu_i h_u l_i + \lambda_{ijk} e_i^c l_j l_k + \lambda_{ijk}' q_i d_j^c l_k + \lambda_{ijk}'' u_i^c d_j^c d_k^c$$

• W_{RPV} violates simultaneously L and B number



$$|\lambda' \cdot \lambda''^*| < 10^{-25} \div 10^{-9} \left(\frac{\tilde{m}}{1 \text{ TeV}}\right)^2$$

[Smirnov, Vissani (1996)]

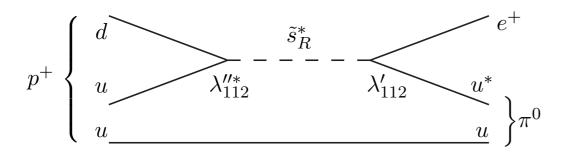


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[Smirnov, Vissani (1996)]

- Possible solution: $M_P = (-)^{3(B-L)}$ or (equivalently) $R_P = M_P (-)^{2S}$
 - B and L accidental global symmetries as in SM
 - LSP is stable (DM candidate)
 - LSP escapes the detector (missing energy)



- R-parity is not necessary
 - Small RPV couplings are ''technically natural'' (holomorphicity of W)
- R-parity is not sufficient for matter stability

$$W_5 \ni \frac{q \, q \, q \, l}{\Lambda} + \frac{u^c u^c d^c e^c}{\Lambda}$$

- R-parity violation might be welcome
 - Neutrino masses within the MSSM field content
 - DM can still be an unstable gravitino
 - Avoids missing energy signatures: LHC bounds can be relaxed



However ...

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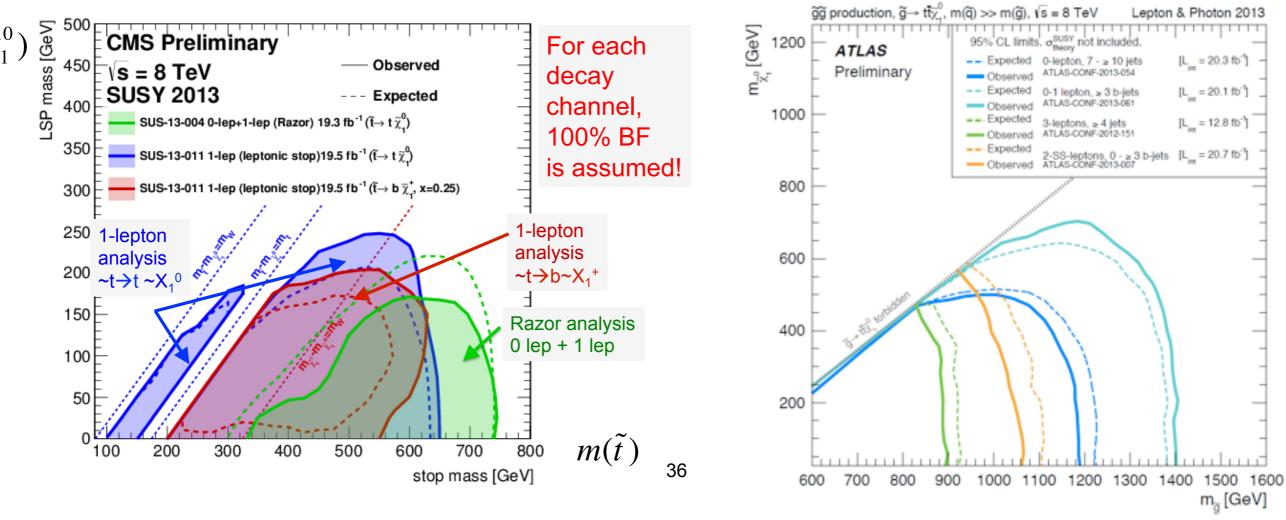
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[...] [Brust, Katz, Lawrence, Sundrum (2012)] [Csaki, Grossman, Heidenreich (2012)] [Graham, Kaplan, Rajendran, Saraswat (2012)] [Allanach, Gripaios (2012)] [Dreiner, Staub, Vicente, Porod (2012)] [Brust, Katz, Sundrum (2012)] [Ruderman, Slayter, Weiner (2012)] [Evans, Kats (2012)] [Asano, Rolbiecki, Sakurai (2013)] [Han, Katz, Son, Tweedie (2013)] [Franceschini, Torre (2013)] [Krnjaic, Stolarski (2013)] [Bhattacherjee, Evans, Ibe, Matsumoto, Yanagida (2013)] [Franceschini, Mohapatra (2013)] [Csaki, Heidenreich (2013)] [Berger, Perelstein, Saelim, Tanedo (2013)] [Florez, Restrepo, Velasquez, Zapata (2013)] [Krnjaic, Tsai (2013)] [Monteux (2013)] [Durieux, Smith (2013)] [...]

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Natural SUSY exclusions $pp \rightarrow \tilde{t}_1 \overline{\tilde{t}_1}$



[See plenary talk by J. Richman and parallel talk by M. D'Alfonso]

 $m_{\tilde{t}} \gtrsim 700 \,\, {
m GeV}$

(barring light-stop window)

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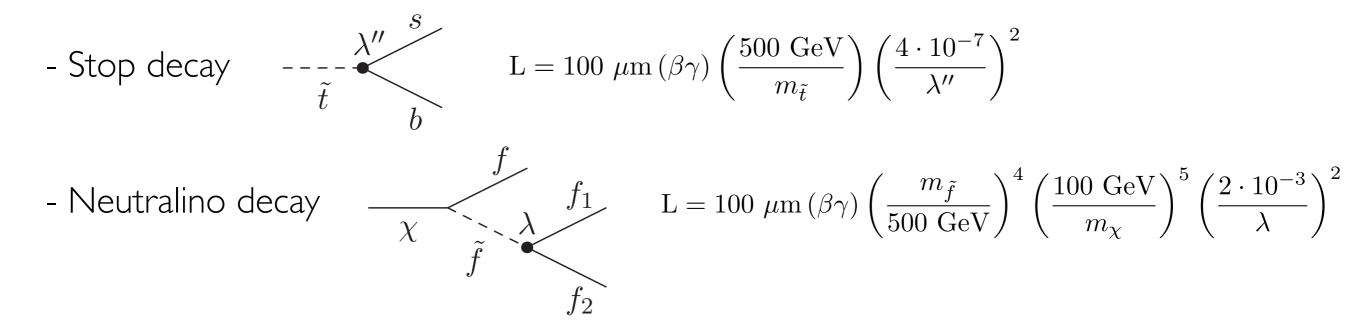
 $m_{\tilde{g}} \gtrsim 1.4 \text{ TeV}$

[See parallel talk by M. Barisonzi]

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R-parity violation @ LHC

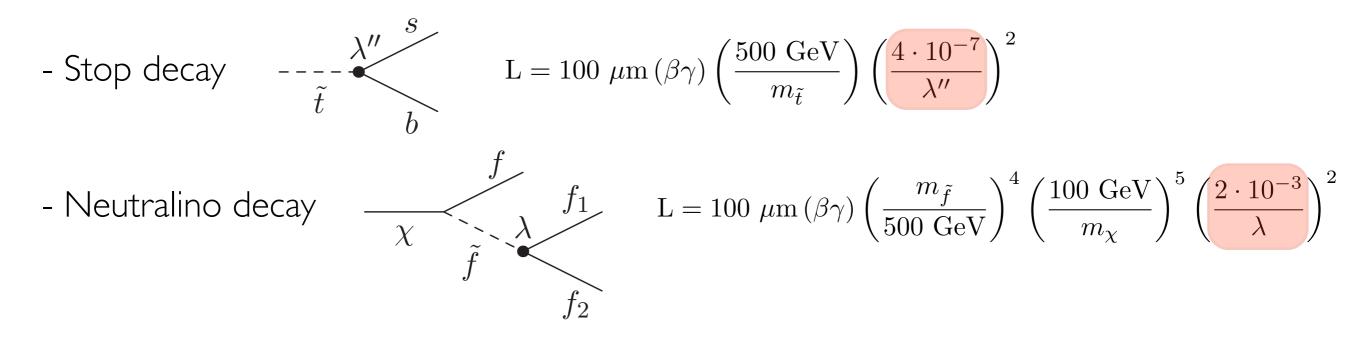
• R-parity violating decays end up into SM fermions





R-parity violation @ LHC

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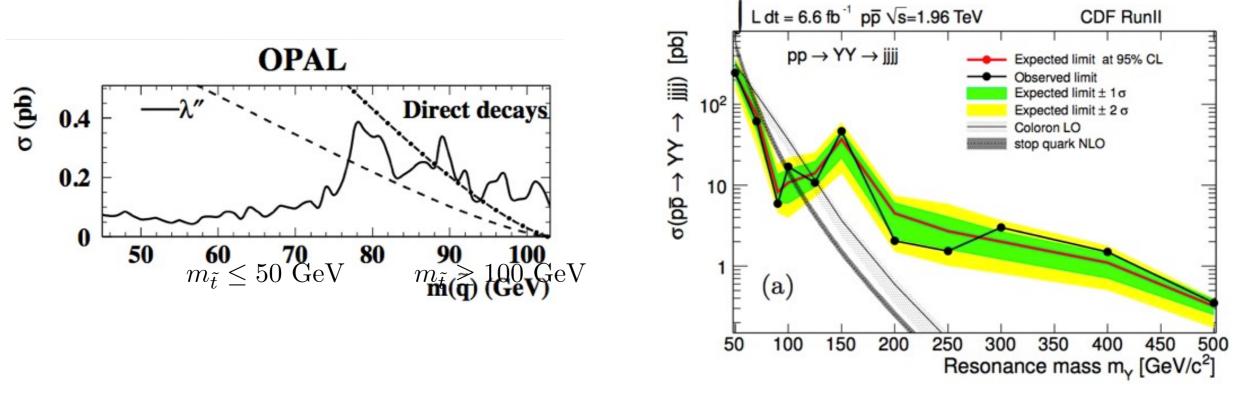


- Prompt decays require RPV couplings of (at least) $\mathcal{O}(10^{-7})$
- Either baryonic or leptonic RPV because of p-decay
 - Leptonic: many leptons in the final state [See however parallel talk by P. Saraswat]
 - Baryonic: better to hide SUSY into QCD backgrounds

Baryonic RPV $sto_{n_{\tilde{t}}(\theta_{\tilde{t}} \le 0.98) \ge 77 \text{ GeV}} m_{\tilde{t}}(\theta_{\tilde{t}} = 0) \ge 88 \text{ GeV}}$

Current limits: LEP + Tevatron

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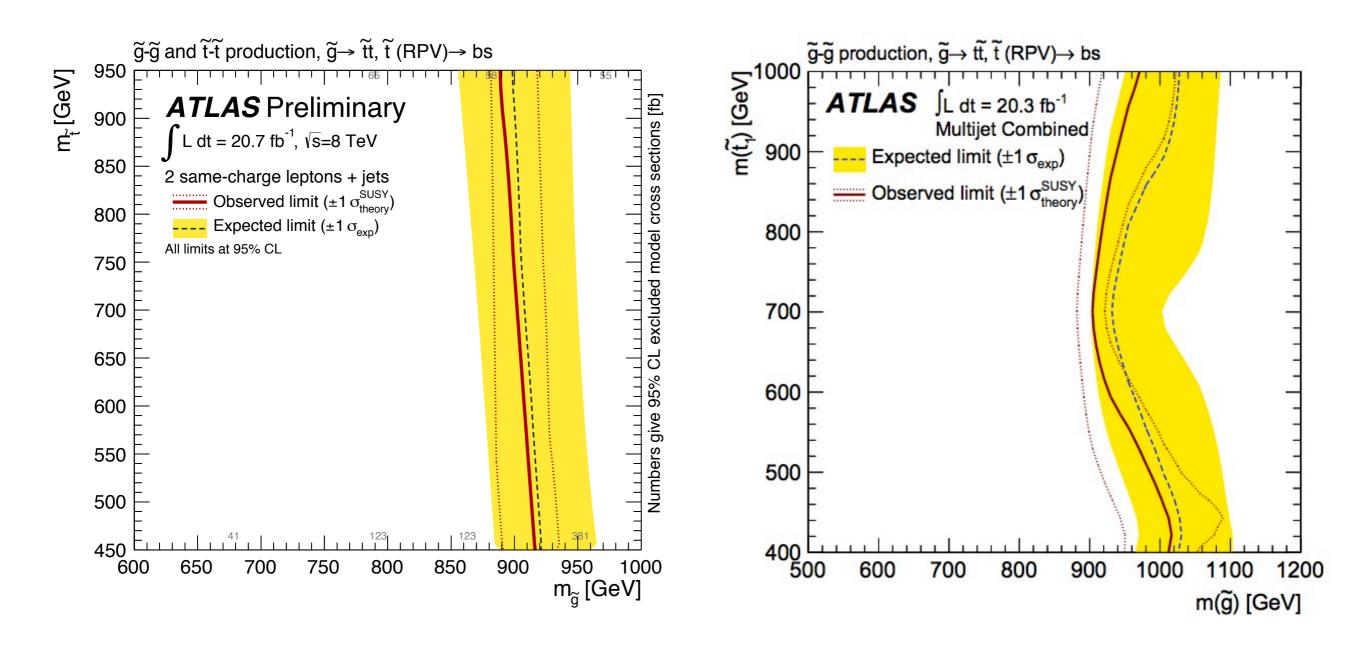


$m_{\tilde{t}} \gtrsim 100 { m ~GeV}$

• LHC not sensitive yet (however b-tagging techniques can improve on that) $m_{\tilde{t}} \leq 50 \text{ GeV}$ $m_{\tilde{t}} \geq 100 \text{ GeV}$ [Franceschini, Torre (2012)]



Baryonic RPV gluinos



 $m_{\tilde{g}} \gtrsim 1 \text{ TeV}$

Summary exclusions

• R-parity conserving

 $m_{\tilde{t}} \gtrsim 700 \text{ GeV}$ $m_{\tilde{g}} \gtrsim 1.4 \text{ TeV}$

Baryonic R-parity violation

 $m_{\tilde{t}} \gtrsim 100 \text{ GeV}$

 $m_{\tilde{g}} \gtrsim 1 \text{ TeV}$

- We improve on naturalness [See however parallel talk by M. Baryakhtar]
- But what about unification ???



RPV & GUT

- Natural expectation: RPV couplings either absent or simultaneously present
- $M_P = (-)^{3(B-L)}$ remnant of gauged B-L ?
- [Martin (1992)] [Mohapatra (1996)] [Aulakh, Bajc, Melfo, Rasin, Senjanovic (2000)] [...]
- Otherwise exact SU(5) invariance $\implies \lambda = \frac{1}{2}\lambda' = \lambda'' \equiv \Lambda$

$$\Lambda_{ijk}\,\overline{5}_i\overline{5}_j10_k\supset\Lambda_{ijk}\,(e_i^cl_jl_k+2\,q_id_j^cl_k+u_i^cd_j^cd_k^c)$$

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- Matter stability requires (at least) $\Lambda_{ijk} < 10^{-10} \left(\frac{\tilde{m}}{1 \text{ TeV}}\right)^2 \ll 10^{-7}$ [Smirnov, Vissani (1996)]
- LSP practically stable on the scale of the detector size
- Sizable RPV couplings apparently an issue for unification !

Baryonic RPV & GUT

- $W_{\rm ren} = W_{\rm MSSM} + \lambda''_{ijk} u^c_i d^c_j d^c_k$ [For soft baryonic RPV see parallel talk by Y.Tsai]
 - Minimal SU(5) + 2 x (5 + 5bar) + fine tuning
 - SU(5) + large representations
 - Flipped SU(5)
 - $SU(5) \times SU(3)$

[Smirnov, Vissani (1996)]

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 - [DL, Nardecchia, Romanino (2013)]
- 4D fully unified gauge group: e.g. SU(5), SO(10), ...
- Small representations: perturbativity up Planck scale
- Renormalizable origin of the extra term
- Absence of fine tuning



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Intuitive idea

• Let us consider SO(10) with the standard embedding of matter

$$16_a = q_a \oplus u_a^c \oplus d_a^c \oplus l_a \oplus e_a^c \oplus \nu_a^c \qquad (a = 1, 2, 3)$$

• To get $u^c d^c d^c$ we need a trilinear term in 16_a

$$\frac{16_H 16_a 16_b 16_c}{\Lambda} \xrightarrow{\langle 16_H \rangle = V_{16}} \frac{V_{16}}{\Lambda} 10_{16_a} \overline{5}_{16_b} \overline{5}_{16_c} \ni e^c l \, l, \, q \, d^c l, \, u^c d^c d^c$$



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- The unwanted operators feature $SU(2)_L$ doublets (q and I)
- An adjoint VEV in the T3R direction can project out the $SU(2)_L$ components

$$\langle 45_H \rangle = V_{45} T_{3R} \qquad (\langle 45_H \rangle 16_a)_{16} = V_{45} (-u_a^c + d_a^c - \nu_a^c + e_a^c)$$

• Analogy with the Dimopoulos-Wilczek mechanism for 2-3 splitting in SO(10)

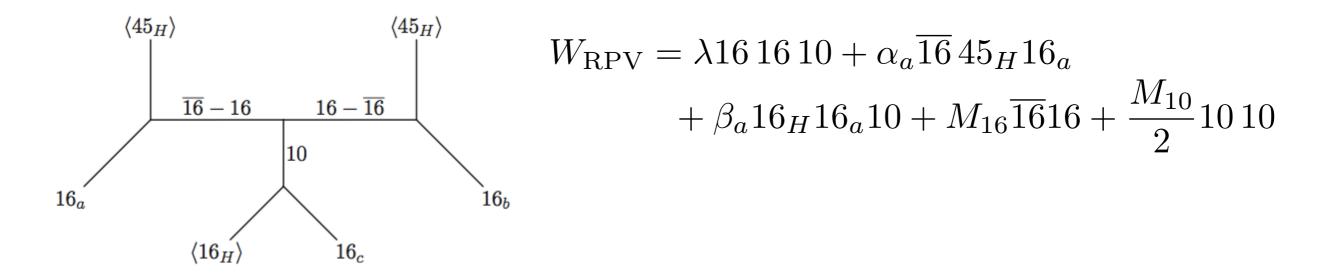
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Intuitive idea

• Hence after SO(10) breaking

$$\frac{16_H (45_H 16_a)_{16} (45_H 16_b)_{16} 16_c}{\Lambda^3} \ni u_a^c d_b^c d_c^c$$

• NR operator points towards its UV completion (not unique)



• Integrating out the vector-like states in the decoupling limit $V_{45}, V_{16} \ll M_{16}, M_{45}$

$$\lambda_{abc}'' = \frac{V_{45}^2 \, V_{16}}{M_{16}^2 \, M_{10}} \, \lambda \, \alpha_a \alpha_{[b} \beta_{c]}$$

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In general ...

• Away from the dec. limit one has to inspect the mass matrices

$$W_{\rm RPV} = \lambda 16\,16\,10 + \alpha_a \overline{16}\,45_H 16_a + \beta_a 16_H 16_a 10 + M_{16} \overline{16} 16 + \frac{M_{10}}{2} 10\,10$$

$$(\overline{d}_{16}^{c} \ \overline{d}_{10}^{c}) \begin{pmatrix} V_{45} \alpha_{a} & M_{16} & 0 \\ V_{16} \beta_{a} & 0 & M_{10} \end{pmatrix} \begin{pmatrix} d_{16}^{c} \\ d_{16}^{c} \\ d_{10}^{c} \end{pmatrix}$$

$$(\overline{l}_{\overline{16}} \ \overline{l}_{10}) \begin{pmatrix} 0 & M_{16} & 0 \\ V_{16} \beta_{a} & 0 & M_{10} \end{pmatrix} \begin{pmatrix} l_{16a} \\ l_{16} \\ l_{10} \end{pmatrix}$$

$$(\overline{u}_{\overline{16}}^{c}) (-V_{45} \alpha_{a} & M_{16}) \begin{pmatrix} u_{16a}^{c} \\ u_{16}^{c} \end{pmatrix}$$

$$(\overline{e}_{\overline{16}}^{c}) (V_{45} \alpha_{a} & M_{16}) \begin{pmatrix} e_{16a}^{c} \\ e_{16}^{c} \end{pmatrix}$$

$$(\overline{q}_{\overline{16}}) (0 & M_{16}) \begin{pmatrix} q_{16a} \\ q_{16} \end{pmatrix}$$

$$16 \supset d_{\text{light}}^c + u_{\text{light}}^c + e_{\text{light}}^c$$

$$10 \supset d_{\text{light}}^c + \ell_{\text{light}}$$

In general ...

• Away from the dec. limit one has to inspect the mass matrices

$$W_{\rm RPV} = \lambda 16\,16\,10 + \alpha_a \overline{16}\,45_H 16_a + \beta_a 16_H 16_a 10 + M_{16} \overline{16} 16 + \frac{M_{10}}{2} 10\,10$$

$$\begin{array}{ll} (\overline{d}_{\overline{16}}^{c} \ \overline{d}_{10}^{c}) \begin{pmatrix} V_{45} \alpha_{a} & M_{16} & 0 \\ V_{16} \beta_{a} & 0 & M_{10} \end{pmatrix} \begin{pmatrix} d_{16a}^{c} \\ d_{10}^{c} \end{pmatrix} & 16 \supset d_{\text{light}}^{c} + u_{\text{light}}^{c} + e_{\text{light}}^{c} \\ (\overline{l}_{\overline{16}} \ \overline{l}_{10}) \begin{pmatrix} 0 & M_{16} & 0 \\ V_{16} \beta_{a} & 0 & M_{10} \end{pmatrix} \begin{pmatrix} l_{16a} \\ l_{16} \\ l_{10} \end{pmatrix} & 10 \supset d_{\text{light}}^{c} + \ell_{\text{light}} \\ (\overline{u}_{\overline{16}}^{c}) (-V_{45} \alpha_{a} & M_{16}) \begin{pmatrix} u_{16a}^{c} \\ u_{16}^{c} \end{pmatrix} & 16 \rightarrow a_{i} u_{i}^{c} \\ (\overline{e}_{\overline{16}}^{c}) (V_{45} \alpha_{a} & M_{16}) \begin{pmatrix} e_{16a}^{c} \\ e_{16}^{c} \end{pmatrix} & 16 \rightarrow b_{i} d_{i}^{c} \\ (\overline{e}_{\overline{16}}^{c}) (V_{45} \alpha_{a} & M_{16}) \begin{pmatrix} e_{16a}^{c} \\ e_{16}^{c} \end{pmatrix} & 10 \rightarrow c_{i} d_{i}^{c} \\ (\overline{q}_{\overline{16}}) (0 & M_{16}) \begin{pmatrix} q_{16a} \\ q_{16} \end{pmatrix} & \lambda 16 \ 16 \ 10 & \Longrightarrow \quad \lambda_{ijk}^{\prime\prime} \propto a_{i} \ (b_{j} c_{k} - b_{k} c_{j}) \end{pmatrix}$$

• Detailed flavour structure depends on the Yukawa sector

Pheno aspects: bounds

- Assume gravitino (LSP) heavier than the proton
- Upper bounds from $\Delta B = 2$ & flavour ($m_{\text{soft}} = \mathcal{O}(500 \text{ GeV})$)

$$\begin{split} |\lambda''_{uds}| &< \mathcal{O}(10^{-5}) \quad [NN \to KK] \,, & |\lambda''_{cdb} \, \lambda''_{csb}| < \mathcal{O}(10^{-3}) \quad [K - \overline{K}] \,, \\ |\lambda''_{udb}| &< \mathcal{O}(10^{-2}) \quad [n - \overline{n}] \,, & |\lambda''_{tdb} \, \lambda''_{tsb}| < \mathcal{O}(10^{-3}) \quad [K - \overline{K}] \,, \\ |\lambda''_{tds}| < \mathcal{O}(10^{-1}) \quad [n - \overline{n}] \,, & |\lambda''_{ids} \, \lambda''_{idb}| < \mathcal{O}(10^{-1}) \quad [B^+ \to K^0 \pi^+] \,, \\ |\lambda''_{tdb}| < \mathcal{O}(10^{-1}) \quad [n - \overline{n}] \,, & |\lambda''_{ids} \, \lambda''_{isb}| < \mathcal{O}(10^{-3}) \quad [B^- \to \phi \pi^-] \,, \end{split}$$

• Lower bounds from prompt decay (e.g. stop NLSP)

$$L = 2 \operatorname{mm} \left(\beta \gamma\right) \left(\frac{500 \text{ GeV}}{m_{\tilde{q}^c}}\right) \left(\frac{0.9 \cdot 10^{-7}}{\lambda''}\right)^2$$

• $10^{-7} \lesssim \lambda'' \lesssim 10^{-5}$ does the job independently of the flavour structure

Pheno aspects: flavour

• GUT correlations

$$\lambda_{ijk}^{\prime\prime} \propto \alpha_i \beta_{[j} \gamma_{k]} \implies \frac{\lambda_{ids}^{\prime\prime}}{\lambda_{jds}^{\prime\prime}} = \frac{\lambda_{idb}^{\prime\prime}}{\lambda_{jdb}^{\prime\prime}} = \frac{\lambda_{isb}^{\prime\prime}}{\lambda_{jsb}^{\prime\prime}} \qquad i, j = u, c, t$$

- relevant when many couplings are at play
- Hierarchical flavour pattern is predicted under assumptions

Yukawa
$$\iff$$
 flavour breaking \implies RPV

- Natural to expect λ_{tbs}'' as the dominant coupling





- The Naturalness status of the MSSM is under pressure
- R-parity violation (especially baryonic) can help
- Requires a quark-lepton asymmetry in apparent contrast with GUT
- Not necessarily ...
- Simple SO(10) models w/o fine-tuning can be conceived



Backup slides





• Add an extra 10_H

 $W_Y = y_{ab} 16_a \ 16_b \ 10_H + y_a 16_a 16 \ 10_H + y \ 16 \ 16 \ 10_H$

- Possibility to fit charged fermions
- Neutrino masses
 - cannot use R-parity violation because of p-decay
 - Δ L=2 effective operator: e.g. $16_a 16_b \overline{16}_H \overline{16}_H / \Lambda$



Yukawa sector

• Connect the flavour breaking in the Yukawa and RPV sectors

$$W_Y = y_{ab} 16_a \ 16_b \ 10_H + y_a 16_a \ 16 \ 10_H + y \ 16 \ 16 \ 10_H$$
$$W_{\rm RPV} = \lambda 16 \ 16 \ 10 + \alpha_a \overline{16} \ 45_H \ 16_a + \beta_a \ 16_H \ 16_a \ 10 + M_{16} \overline{16} \ 16 + \frac{M_{10}}{2} \ 10 \ 10$$

- Horizontal SU(3)_H symmetry (16_a ~ triplet)
- SU(3)_H broken by two hierarchical spurions A >> B (A and B ~ anti-triplet)

$$\begin{array}{l} \alpha_{a} = r_{\alpha}A_{a} + s_{\alpha}B_{a} , & \epsilon \ll 1 \\ \beta_{a} = r_{\beta}A_{a} + s_{\beta}B_{a} , & \Longleftrightarrow \\ y_{a} = r_{z}A_{a} + s_{z}B_{a} , & & \swarrow \\ y_{ab} = r_{y}A_{a}A_{b} + s_{y}B_{a}B_{b} + t_{y} \left(A_{a}B_{b} + B_{a}A_{b}\right) , \end{array}$$

$$(\alpha_a) = \alpha(0, 0, 1)$$

$$(\beta_a) = \beta(0, \epsilon, 1)$$

$$y_{33} \sim y_3 = \mathcal{O}(1)$$

$$y_{23} = y_{32} \sim y_2 = \mathcal{O}(\epsilon)$$

$$y_{22} = \mathcal{O}(\epsilon^2)$$

- Predicts hierarchical pattern $\lambda_{cbs}^{\prime\prime} = -\epsilon \frac{y_{23}}{y_{33}} \lambda_{tbs}^{\prime\prime}$

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- Qualitatively different from other frameworks like MFV-RPV

PV [Nikolidakis, Smith (2008)] [Csaki, Grossman, Heidenreich (2012)]

VEV alignment

$$W_{\text{vev-align.}} = m_1 45_H^2 + m_2 54_H^2 + \lambda_1 54_H 45_H^2 + \lambda_2 54_H^3$$

$$\langle 45_H \rangle = \operatorname{diag}(V_R, V_R, V_{B-L}, V_{B-L}, V_{B-L}) \otimes i\sigma_2$$

$$\langle 54_H \rangle = \operatorname{diag}(-\frac{3}{2}V_{54}, -\frac{3}{2}V_{54}, V_{54}, V_{54}, V_{54}) \otimes I$$

- SUSY vacuum implies $F_{45_{H}}=0$, namely

 $(m_1 - \frac{3}{2}\lambda_1 V_{54})V_R = 0$ $(m_1 + \lambda_1 V_{54})V_{B-L} = 0$

• A solution is provided by $m_1 = \frac{3}{2}\lambda_1 V_{54}$ and $V_{B-L} = 0$

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